

Study of the ^{19}Ne spectroscopic properties of astrophysical interest via a new method of inelastic scattering
(November 2013)

Florent Boulay (GANIL, France)

 Nova Del 2013

1) Astrophysical motivations
(Nova)

2) The experiment and preliminary results
(Inelastic scattering of ^{19}Ne)

3) Angular distribution

1) *Astrophysical context*

2) *The experiment*

3) *Angular distribution*

Study of the ^{19}Ne spectroscopic properties of astrophysical interest via a new method of inelastic scattering
(November 2013)

 Nova Del 2013

Authentic from Normandie

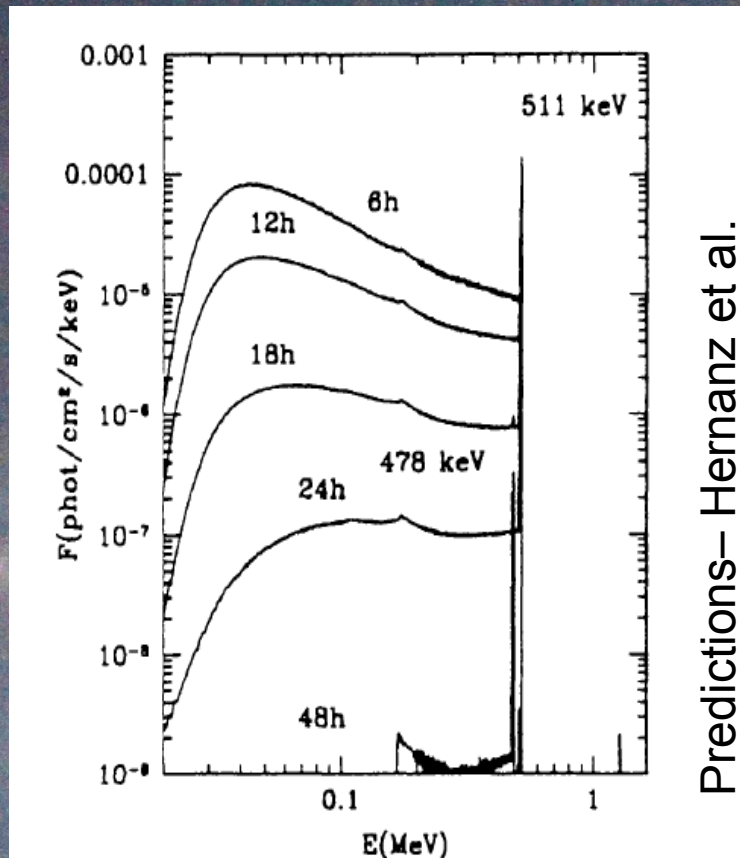


Picturing by François De Oliveira

1/13

Study of the ^{19}Ne spectroscopic properties of astrophysical interest via a new method of inelastic scattering (November 2013)

Nova Del 2013



- A key observable: Gamma rays at 511 keV
- One of the main β^+ emitters: ^{18}F
- 2 main reactions constrain the abundance of ^{18}F : $^{18}\text{F}(p,\alpha)^{15}\text{O}$ et $^{18}\text{F}(p,\gamma)^{19}\text{Ne}$.
- Study of ^{18}F via ^{19}Ne .

b) Knowledge about ^{19}Ne
 (Gamow Window 300 GK)
 

No	E_x^a (MeV)	J^π^b	Γ_γ^c (eV)	Γ_p^d (keV)	Γ_α^e (keV)
1	6.419	$(\frac{3}{2}^+)$	0.77(41)	2.2(4)E-37	0.27(27)
2	(6.422)	$(\frac{11}{2}^+)$	0.35(18)	1.8(18)E-38	20(14)E-3
3	6.437	$\frac{1}{2}^-$	[1(1)]	1.1(11)E-20	220(20) (M)
4	6.449	$(\frac{3}{2}^+)$	1.1(6)	4(4)E-15	1.3(10)
5	(6.504)	$(\frac{7}{2}^+)$	0.14(8)	4.6(46)E-10	0.4(4)
6	(6.542)	$(\frac{9}{2}^+)$	0.30(16)	2.7(27)E-12	1.3(11)E-2
7	6.698	$(\frac{5}{2}^+)$	0.29(15)	1.2(12)E-5	1.2(10)
8	6.741	$\frac{3}{2}^-$	5.0(26)	2.22(69)E-3	5.2(37)
9	(6.841)	$(\frac{3}{2}^-)$	2.8(15)	9.7(97)E-3	25(18)
10	6.861	$\frac{7}{2}^-$	2.3(12)	1.1(11)E-5	1.2(0.9)
11	(6.939)	$(\frac{1}{2}^-)$	[1(1)]	3.4(34)E-2	99(69)
12	(7.054)	$(\frac{5}{2}^+)$	[1(1)]	4.7(47)E-2	29(25)
13	7.0757	$\frac{3}{2}^+$	[1(1)]	15.2(1)	23.8(12) (M)
14	7.173	$(\frac{11}{2}^-)$	0.15(8)	9.8(98)E-8	1.2(10)E-2
15	7.238	$\frac{3}{2}^+$	[1(1)]	0.35(35)	6.0(52)

 Values C. D. Nesaraja *et al.* (2007)

b) Knowledge about ^{19}Ne
 (Gamow Window 300 GK)
 

No	E_x^a (MeV)	—	J^π^b	Γ_γ^c (eV)	Γ_p^d (keV)	Γ_α^e (keV)
1	6.419		$(\frac{3}{2}^+)$	0.77(41)	2.2(4)E-37	0.27(27)
2	(6.422)		$(\frac{11}{2}^+)$	0.35(18)	1.8(18)E-38	20(14)E-3
3	6.437		$\frac{1}{2}^-$	[1(1)]	1.1(11)E-20	220(20) (M)
4	6.449		$(\frac{3}{2}^+)$	1.1(6)	4(4)E-15	1.3(10)
5	(6.504)		$(\frac{7}{2}^+)$	0.14(8)	4.6(46)E-10	0.4(4)
6	(6.542)		$(\frac{9}{2}^+)$	0.30(16)	2.7(27)E-12	1.3(11)E-2
7	6.698		$(\frac{5}{2}^+)$	0.29(15)	1.2(12)E-5	1.2(10)
8	6.741		$\frac{3}{2}^-$	5.0(26)	2.22(69)E-3	5.2(37)
9	(6.841)		$(\frac{3}{2}^-)$	2.8(15)	9.7(97)E-3	25(18)
10	6.861		$\frac{7}{2}^-$	2.3(12)	1.1(11)E-5	1.2(0.9)
11	(6.939)		$(\frac{1}{2}^-)$	[1(1)]	3.4(34)E-2	99(69)
12	(7.054)		$(\frac{5}{2}^+)$	[1(1)]	4.7(47)E-2	29(25)
13	7.0757		$\frac{3}{2}^+$	[1(1)]	15.2(1)	23.8(12) (M)
14	7.173		$(\frac{11}{2}^-)$	0.15(8)	9.8(98)E-8	1.2(10)E-2
15	7.238		$\frac{3}{2}^+$	[1(1)]	0.35(35)	6.0(52)

 Values C. D. Nesaraja *et al.*(2007)

b) Knowledge about ^{19}Ne

A. M. Laird *et al* PRL (2012) : $^{19}\text{F}(^3\text{He},t)^{19}\text{Ne}$.

No	E_x^a (MeV)	J^π^b	Γ_γ^c (eV)	Γ_p^d (keV)	Γ_α^e (keV)
1	6.419	$(\frac{3}{2}^+)$	0.7(41)	2.2(4)E-37	0.27(27)
2	(6.422)	$(\frac{11}{2}^+)$	0.35(18)	1.8(18)E-38	20(14)E-3
3	6.437	$\frac{1}{2}^-$	[1(1)]	1.1(11)E-20	220(20) (M)
4	6.449	$(\frac{3}{2}^+)$	1.1(6)	4(4)E-15	1.3(10)
5	(6.504)	$(\frac{7}{2}^+)$	0.14(8)	4.6(46)E-10	0.4(4)
6	(6.542)	$(\frac{9}{2}^+)$	0.30(16)	2.7(27)E-12	1.3(11)E-2
7	6.698	$(\frac{5}{2}^+)$	0.29(15)	1.2(12)E-5	1.2(10)
8	6.741	$\frac{3}{2}^-$	5.0(26)	2.22(69)E-3	5.2(37)
9	(6.841)	$(\frac{3}{2}^-)$	2.8(15)	9.7(97)E-3	25(18)
10	6.861	$\frac{7}{2}^-$	2.3(12)	1.1(11)E-5	1.2(0.9)
11	(6.939)	$(\frac{1}{2}^-)$	[1(1)]	3.4(34)E-2	99(69)
12	(7.054)	$(\frac{5}{2}^+)$	[1(1)]	4.7(47)E-2	29(25)
13	7.0757	$\frac{3}{2}^+$	[1(1)]	15.2(1)	23.8(12) (M)
14	7.173	$(\frac{11}{2}^-)$	0.15(8)	9.8(98)E-8	1.2(10)E-2
15	7.238	$\frac{3}{2}^+$	[1(1)]	0.35(35)	6.0(52)

Controversery on 3/2 states

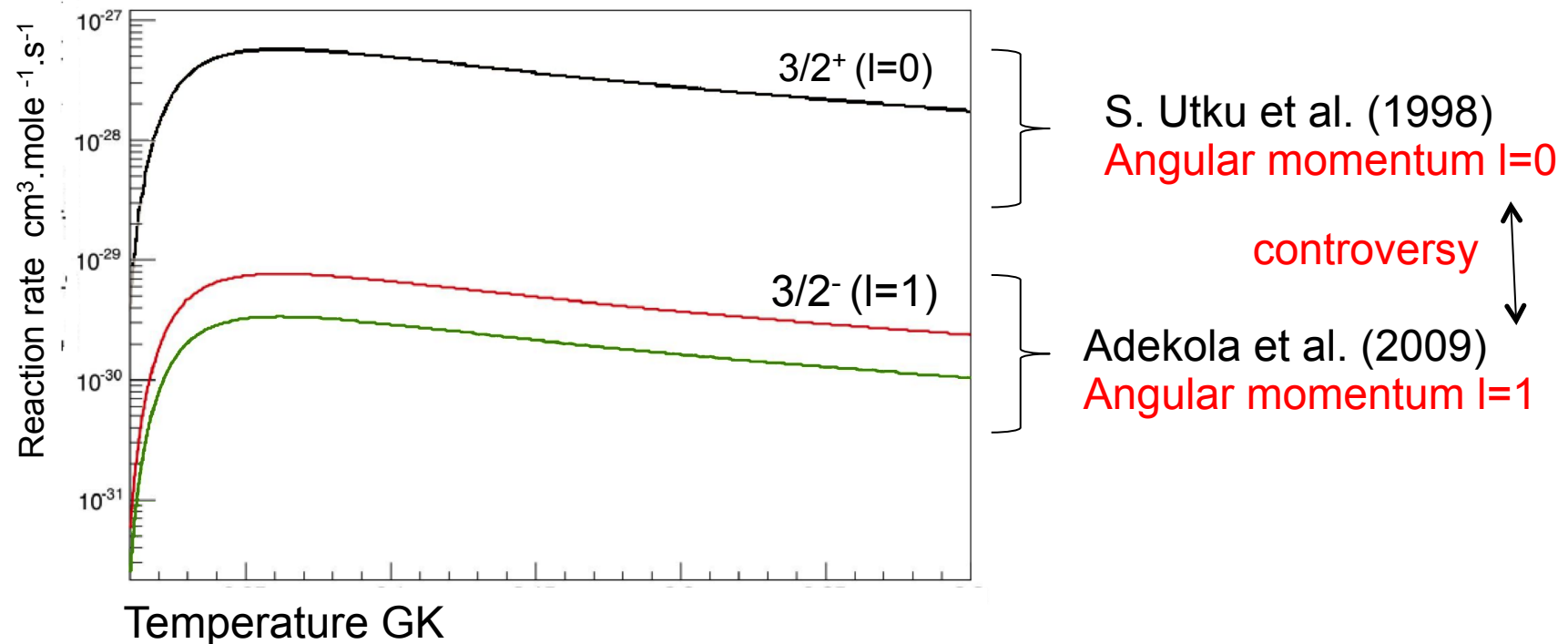
(Gamow Window 300 GK)

D. J. Mountford *et al*. PRC (2012) : $^{18}\text{F}(p,\alpha)^{19}\text{Ne}$ @ 3.924MeV/u

Values C. D. Nesaraja *et al.*(2007)

c) An example : the resonant state at 8 keV

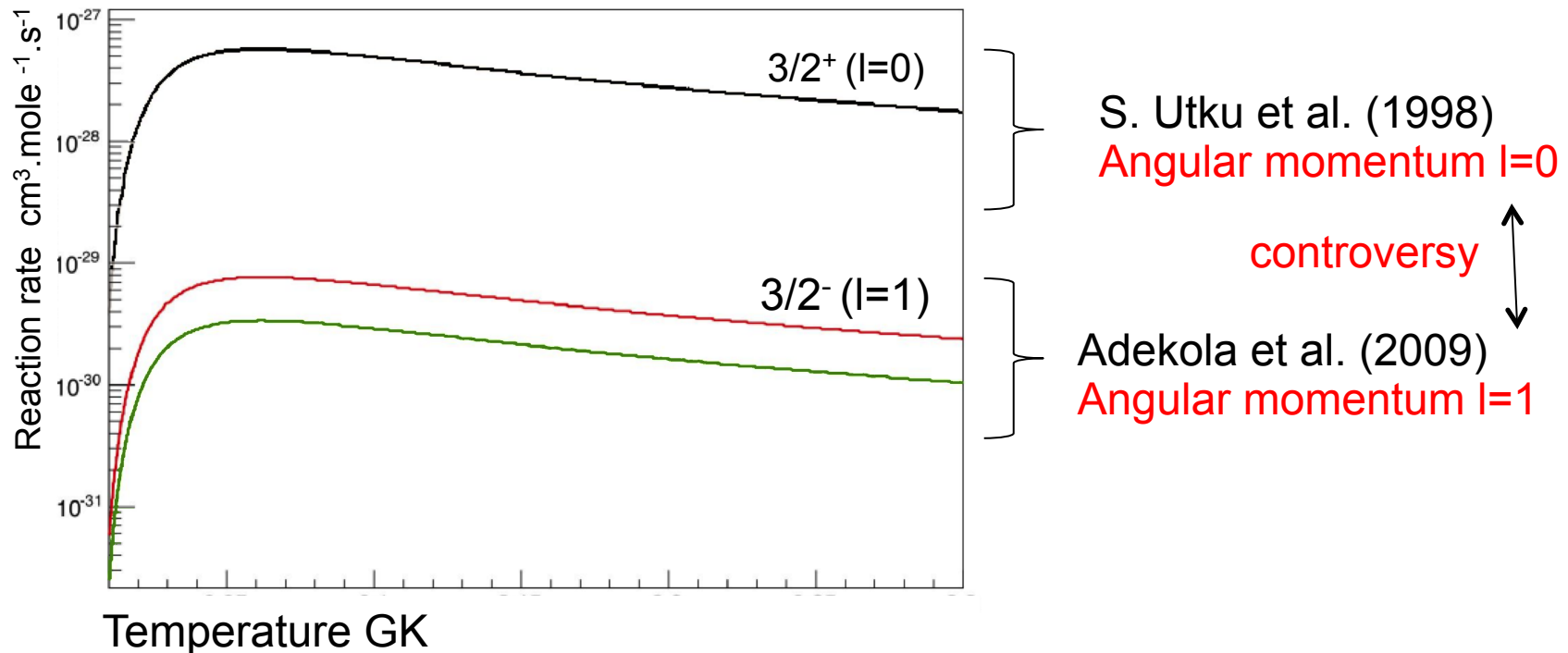
Reaction rate for resonant state at 8 keV of the reaction $^{18}\text{F}(p,\alpha)^{15}\text{O}$



F. Boulay Master Thesis (2012)

c) An example : the resonant state at 8 keV

Reaction rate for resonant state at 8 keV of the reaction $^{18}\text{F}(p,\alpha)^{15}\text{O}$



=> We do need to measure experimentally the spectroscopic properties of ^{19}Ne

The e641s experiment at GANIL

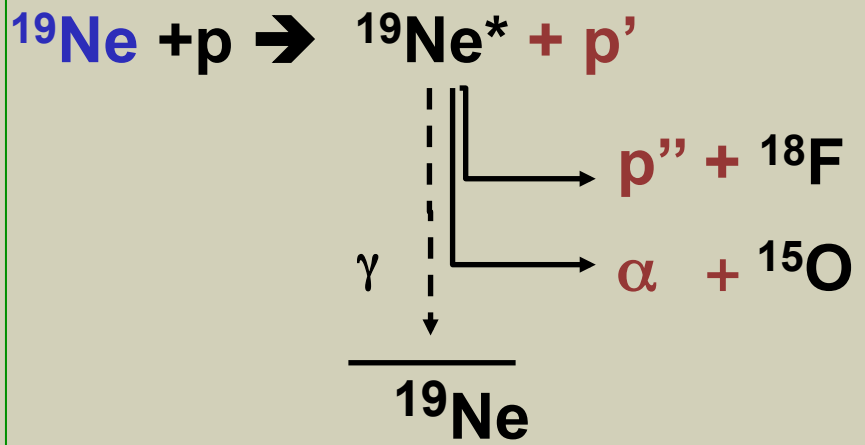
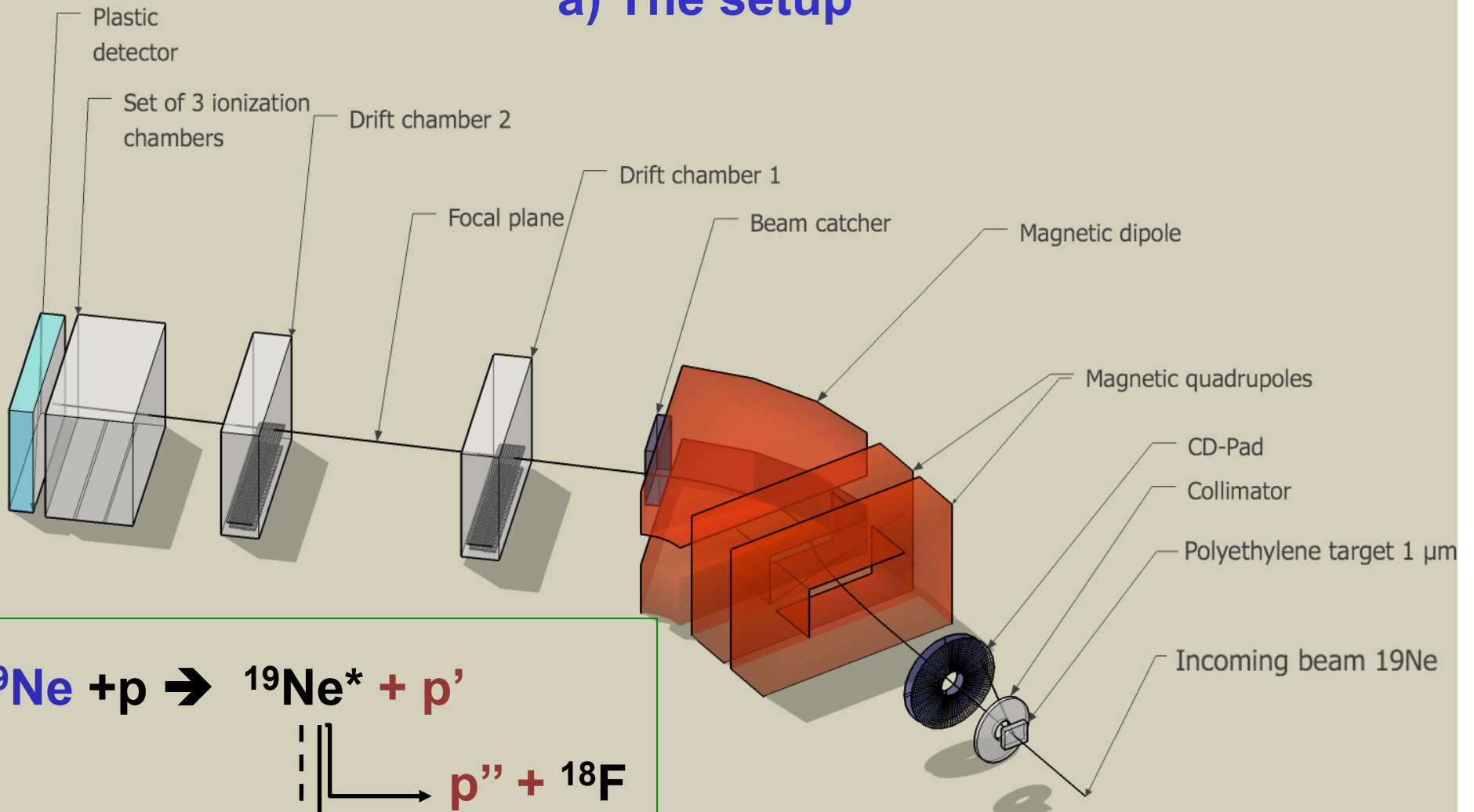
November 2013

1) Astrophysical context

2) The experiment

3) Angular distribution

a) The setup



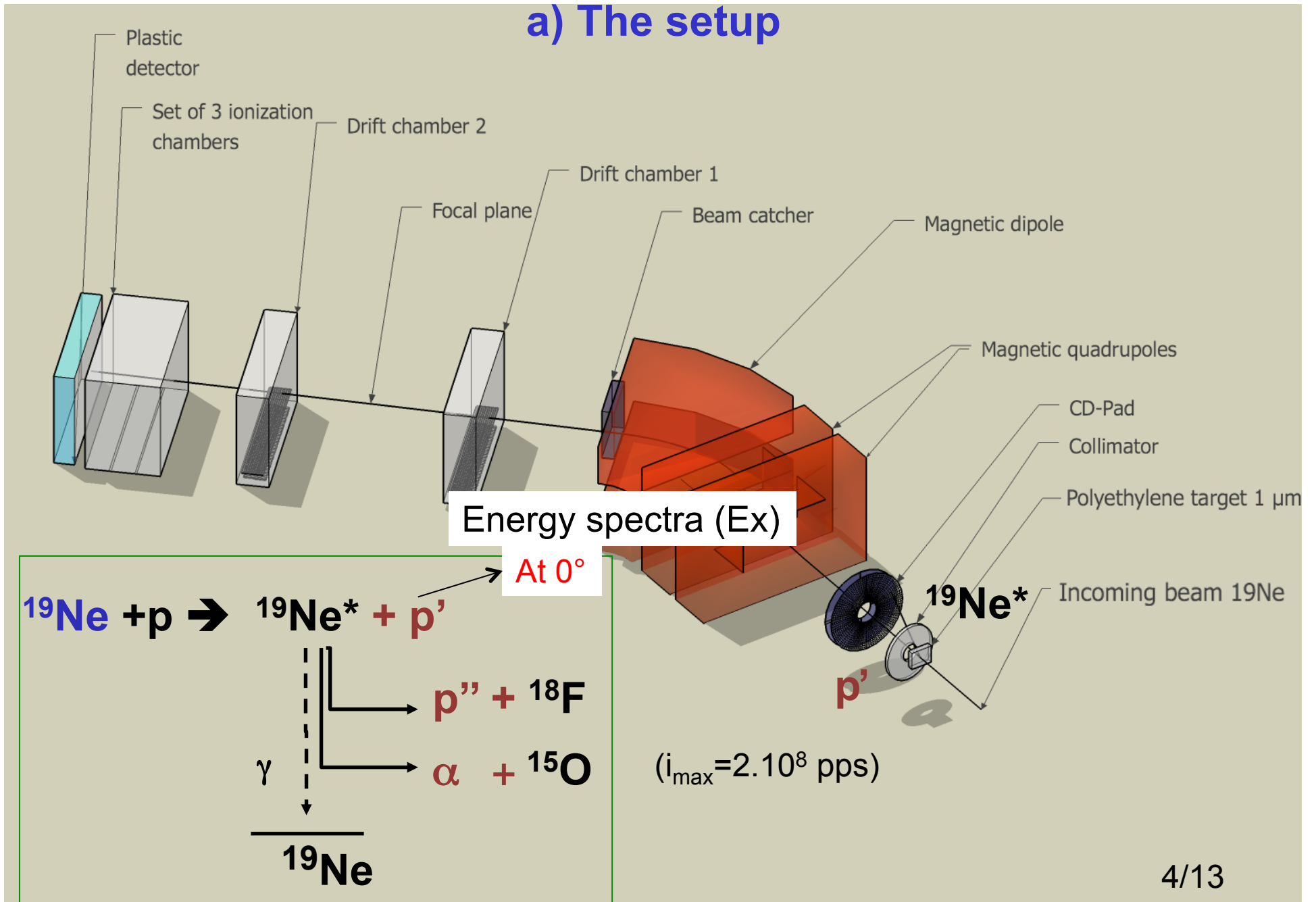
($i_{\text{max}} = 2 \cdot 10^8$ pps)

1) Astrophysical context

2) The experiment

3) Angular distribution

a) The setup



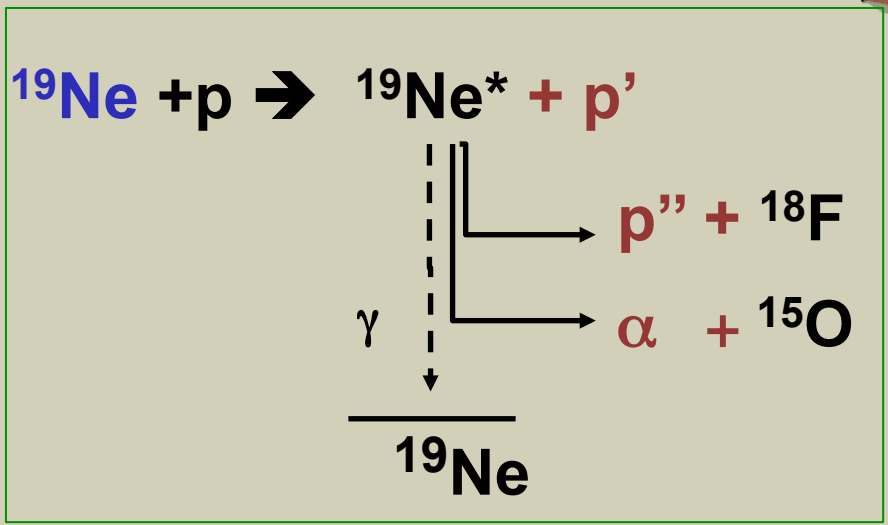
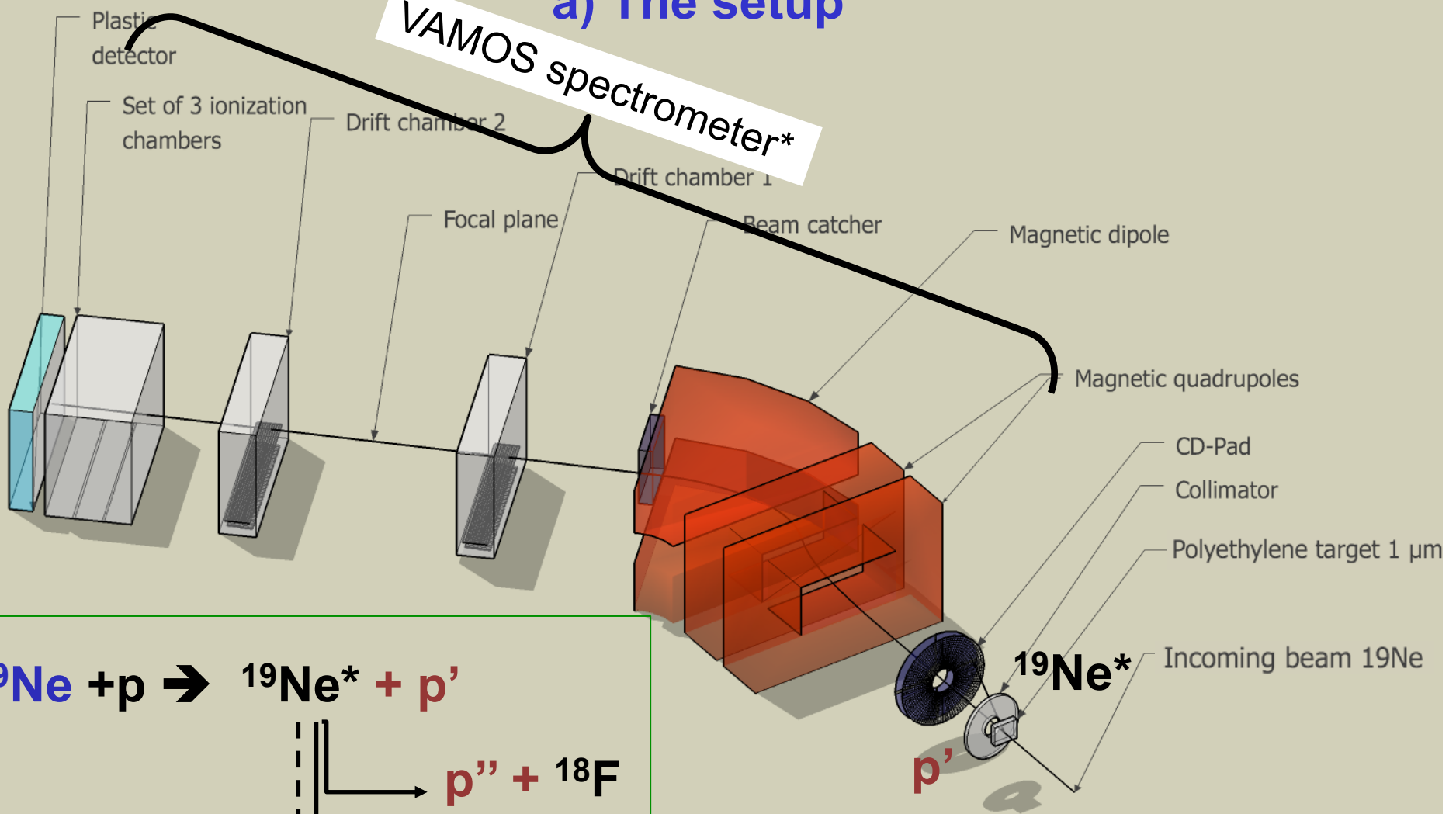
1) Astrophysical context

2) The experiment

3) Angular distribution

a) The setup

VAMOS spectrometer*



($i_{\max} = 2.10^8$ pps)

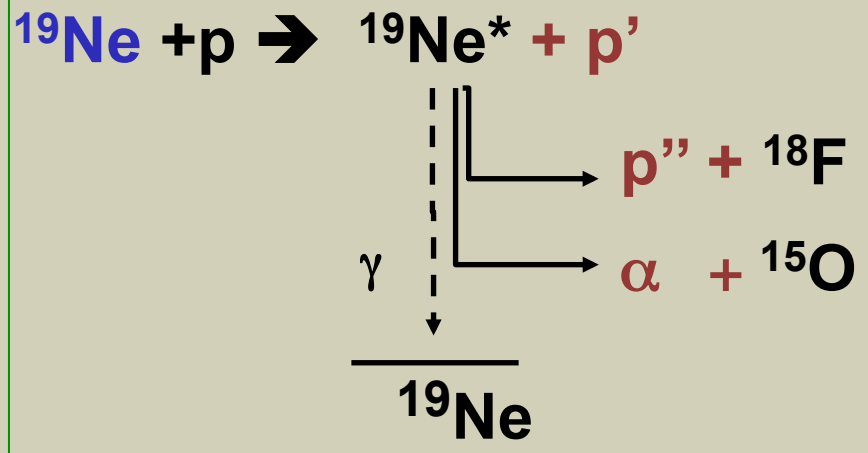
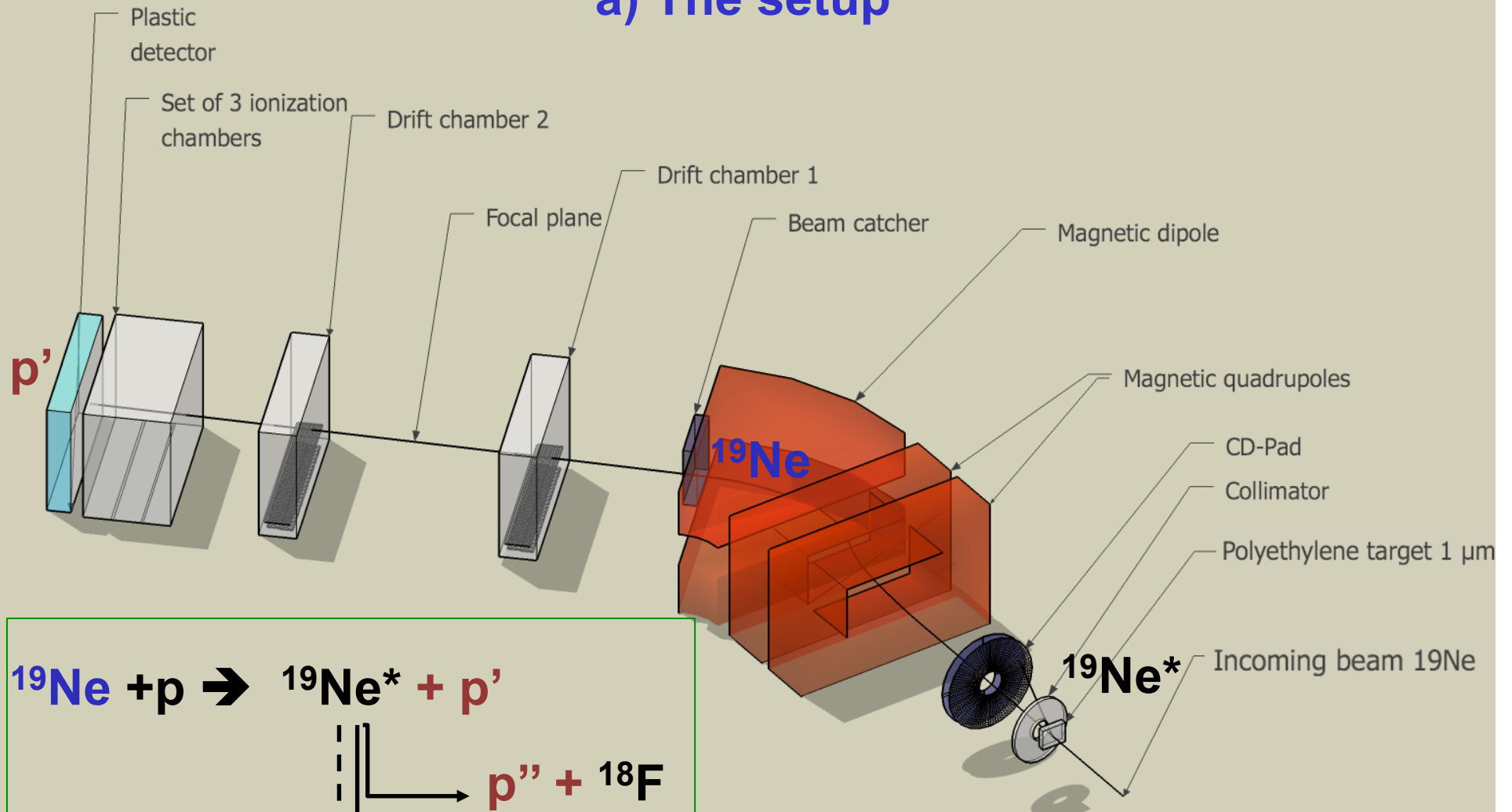
*M. Rejmund *et al.* NIMA (2011) 5/13

1) Astrophysical context

2) The experiment

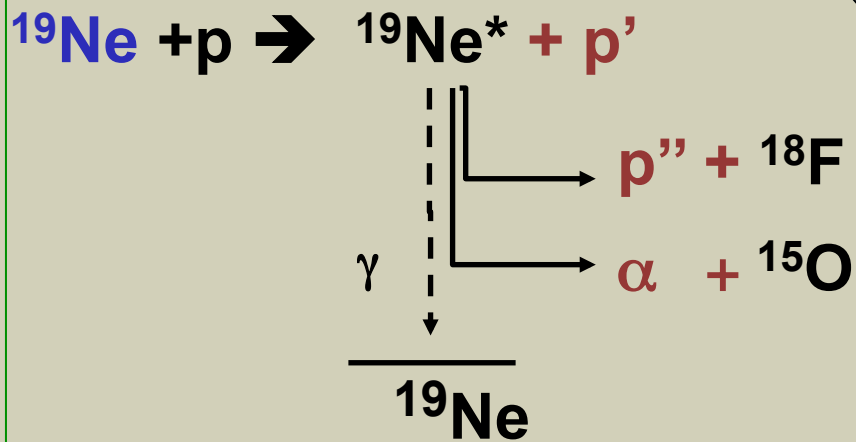
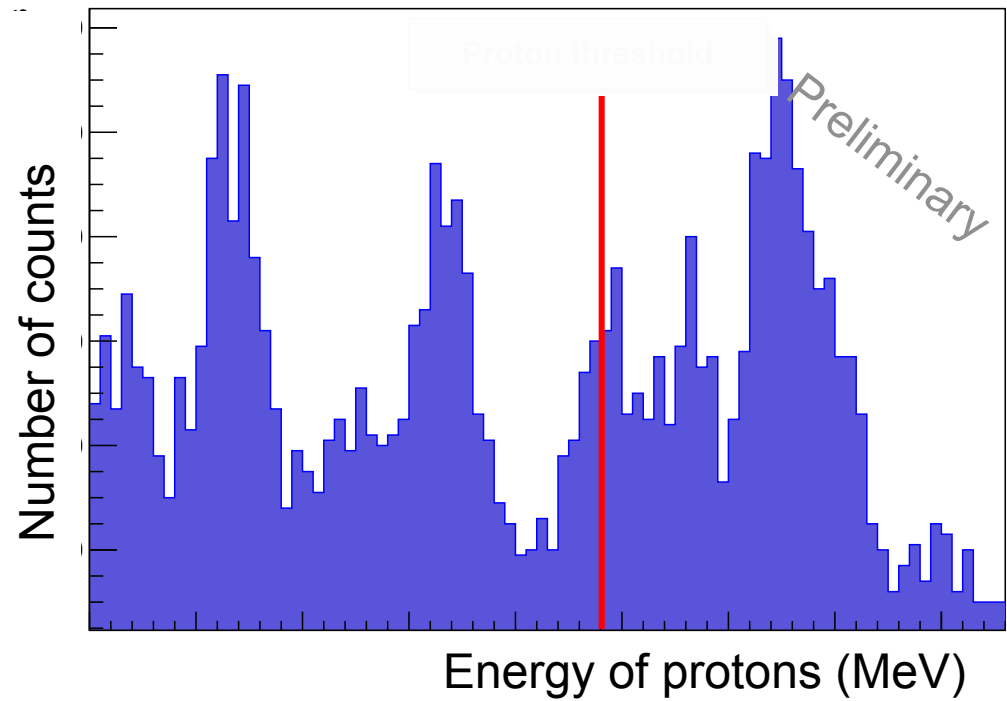
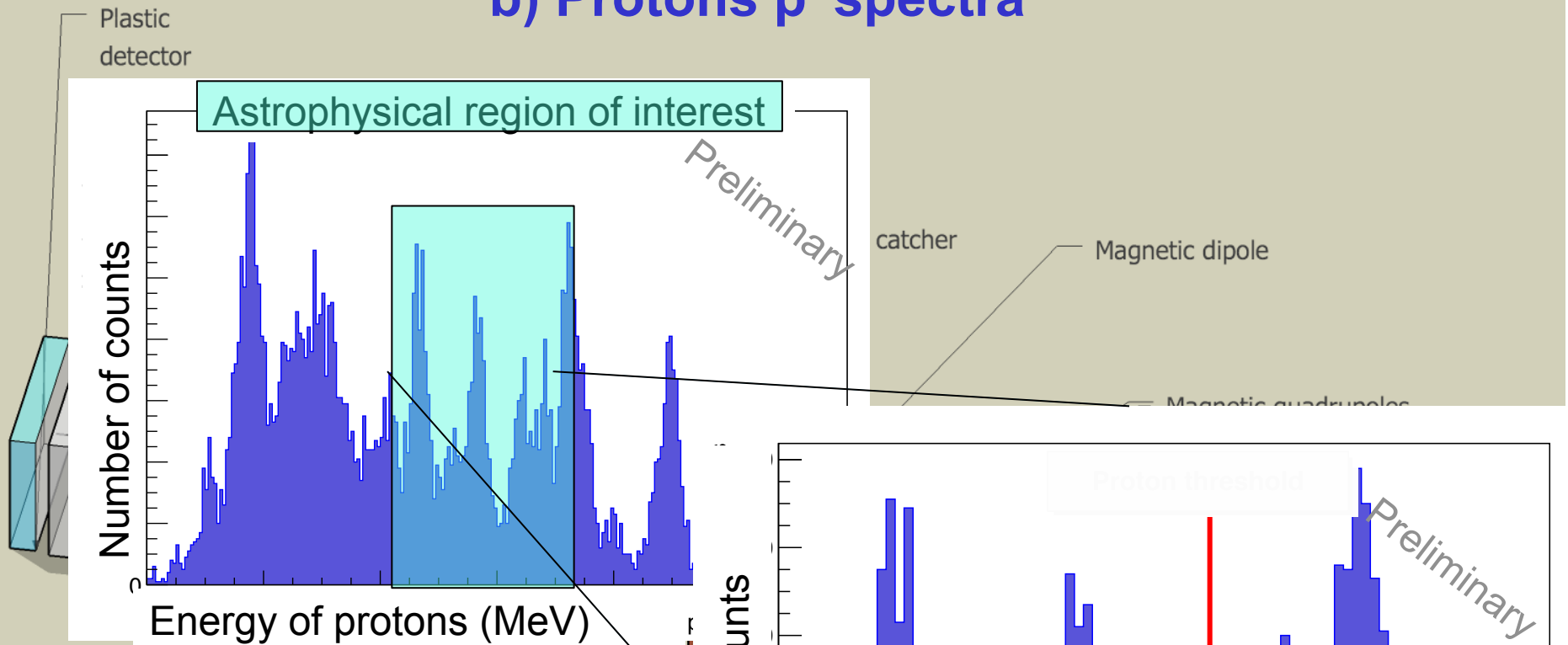
3) Angular distribution

a) The setup



($i_{\text{max}} = 2 \cdot 10^8$ pps)

b) Protons p' spectra



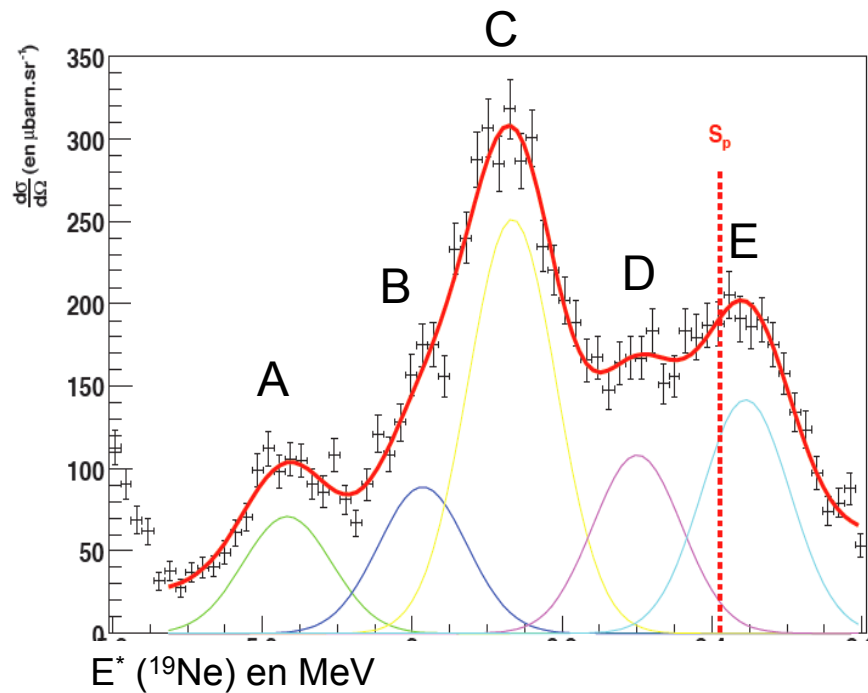
1) Astrophysical context

2) The experiment

3) Angular distribution

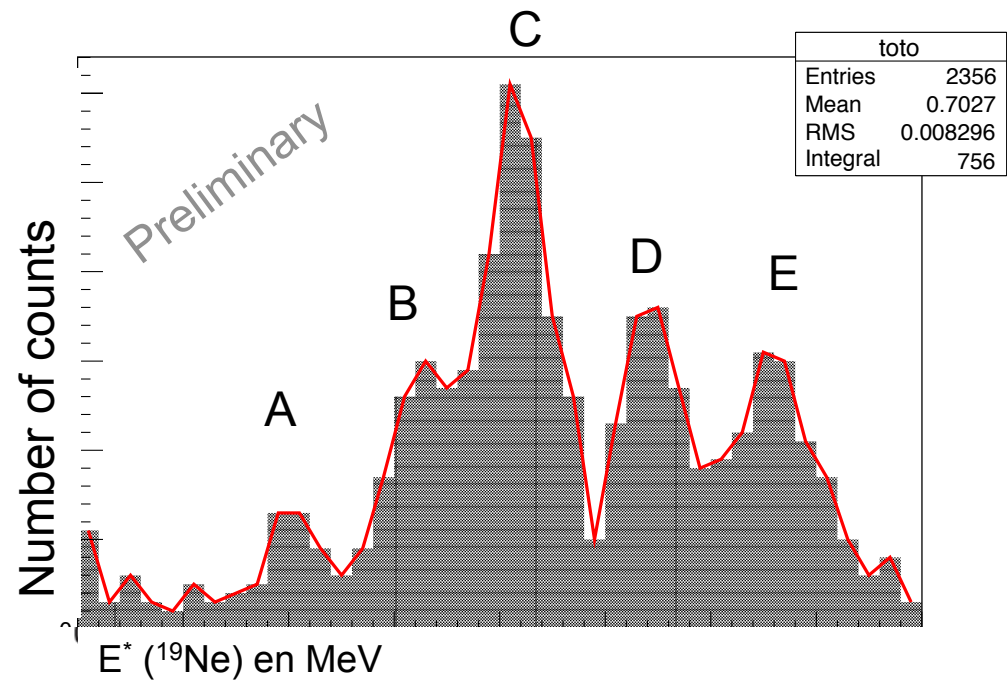
c) $^{19}\text{Ne}^*$ excitation energy spectra

Comparison in the range of excitation energy of $^{19}\text{Ne}^*$ (5.6 MeV and 6.6 MeV)



Ph.D JC Dalouzy

Resolution in excitation energy
140 keV



This experiment

Resolution in excitation energy
around 80 keV

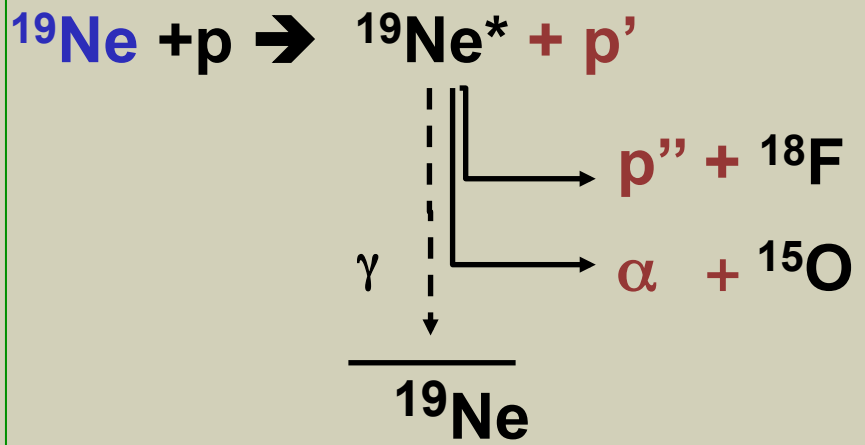
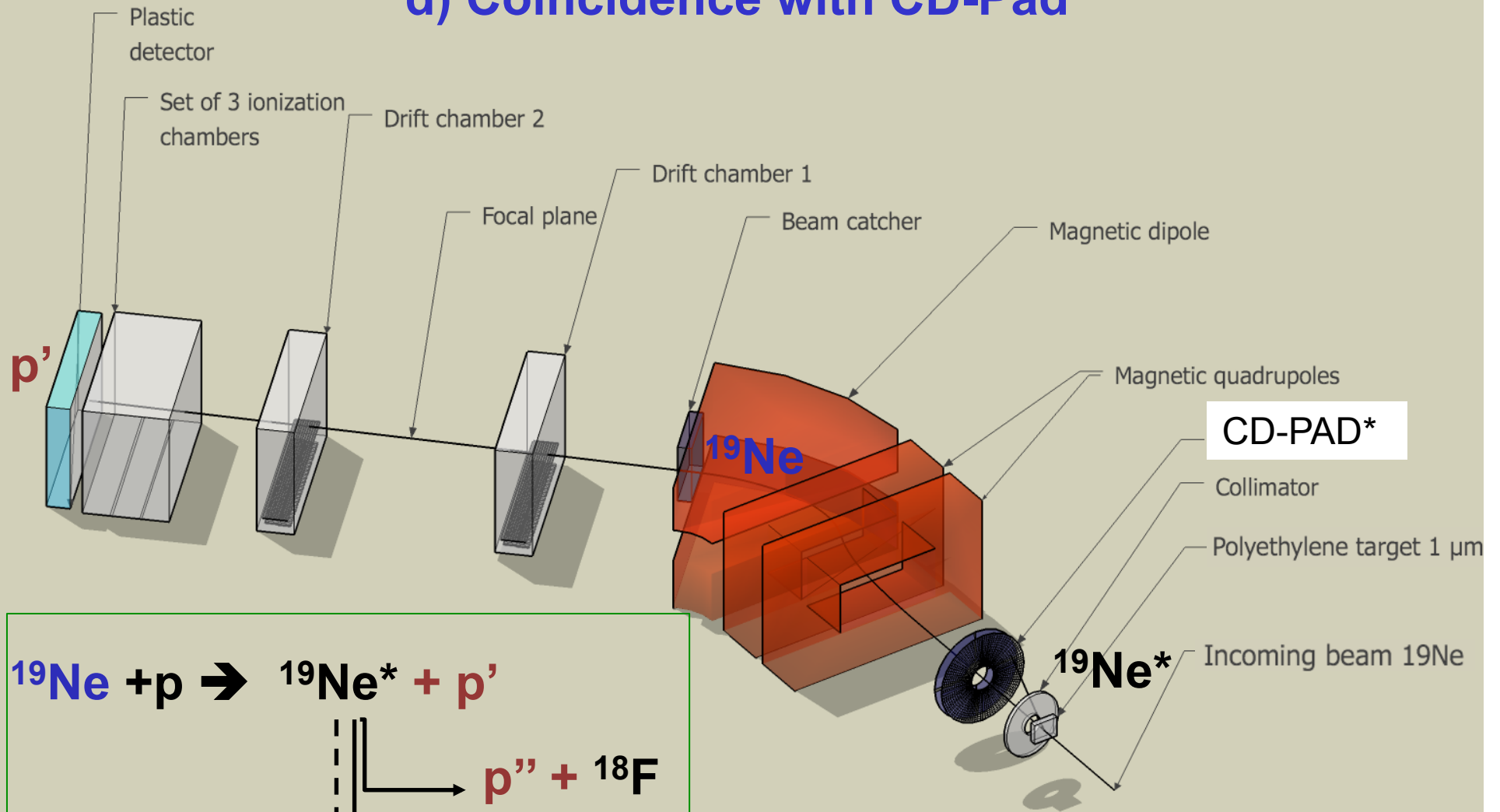
Improvement in resolution by a factor ~ 2 with the new method !!!!!

1) Astrophysical context

2) The experiment

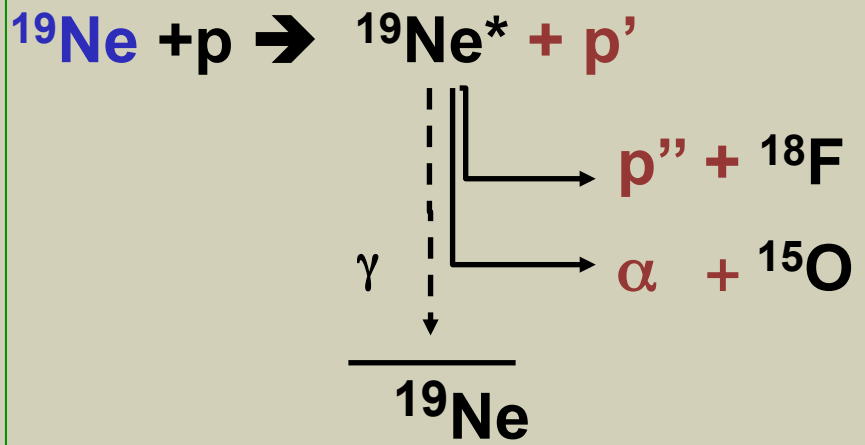
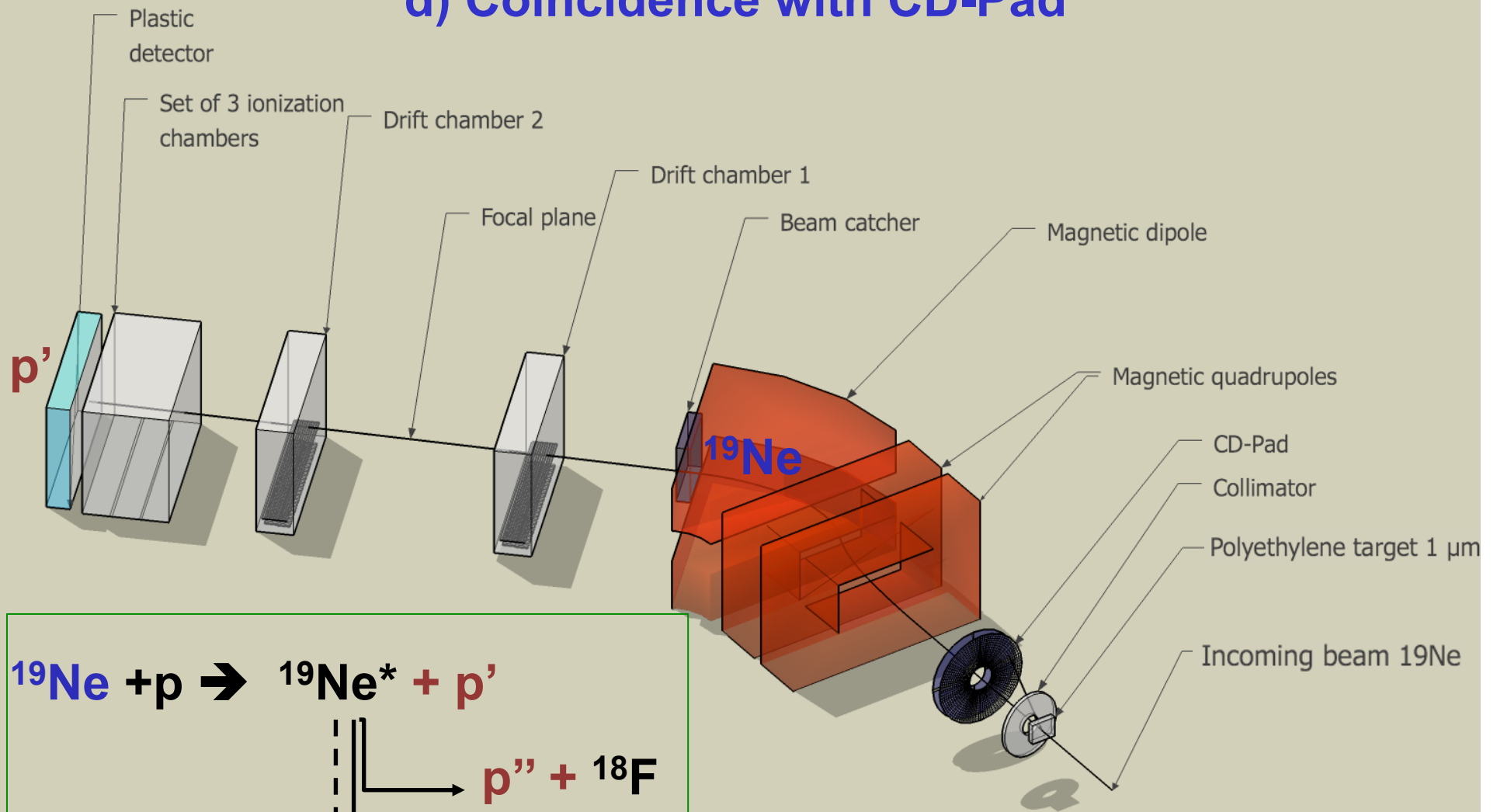
3) Angular distribution

d) Coincidence with CD-Pad

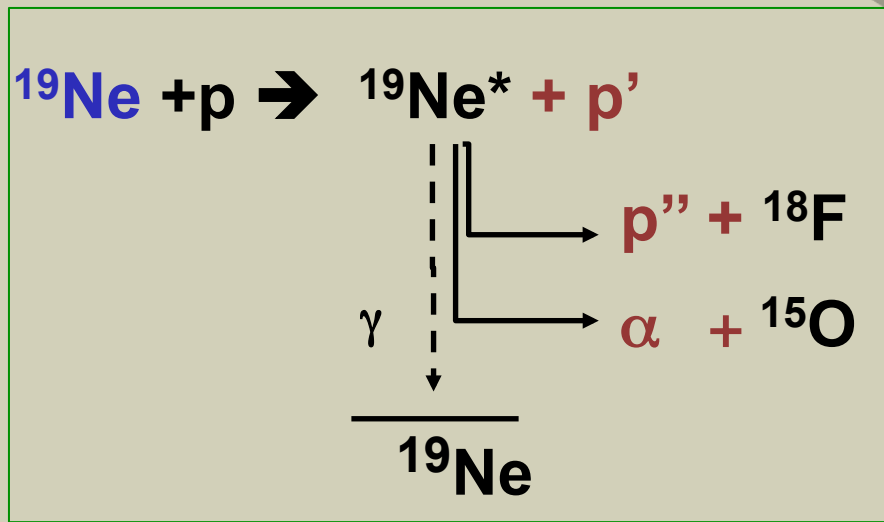
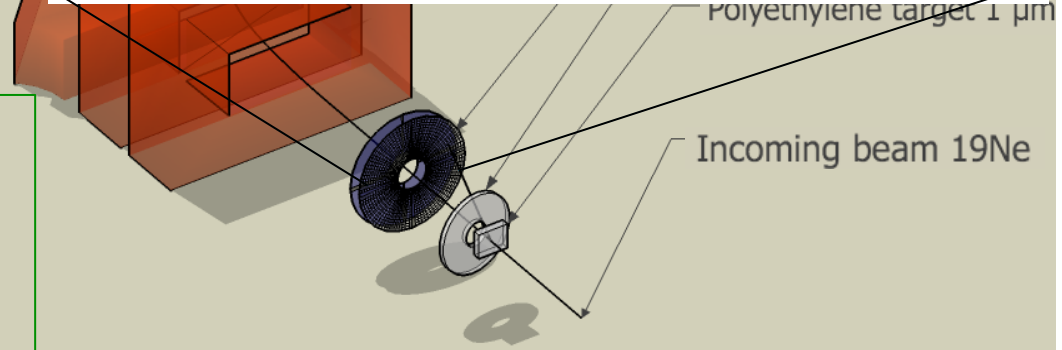
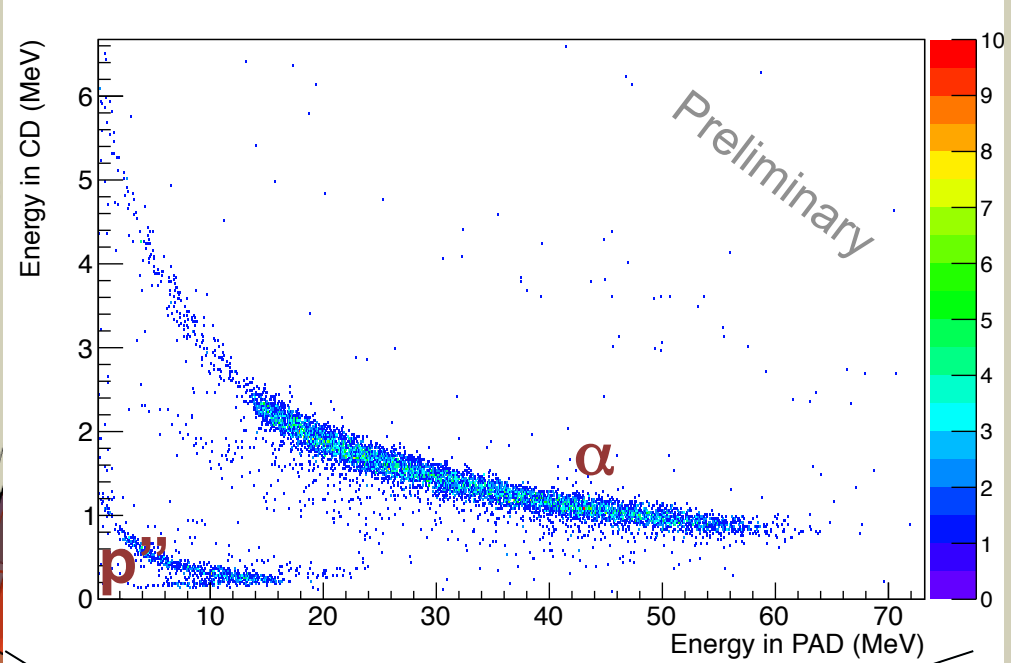
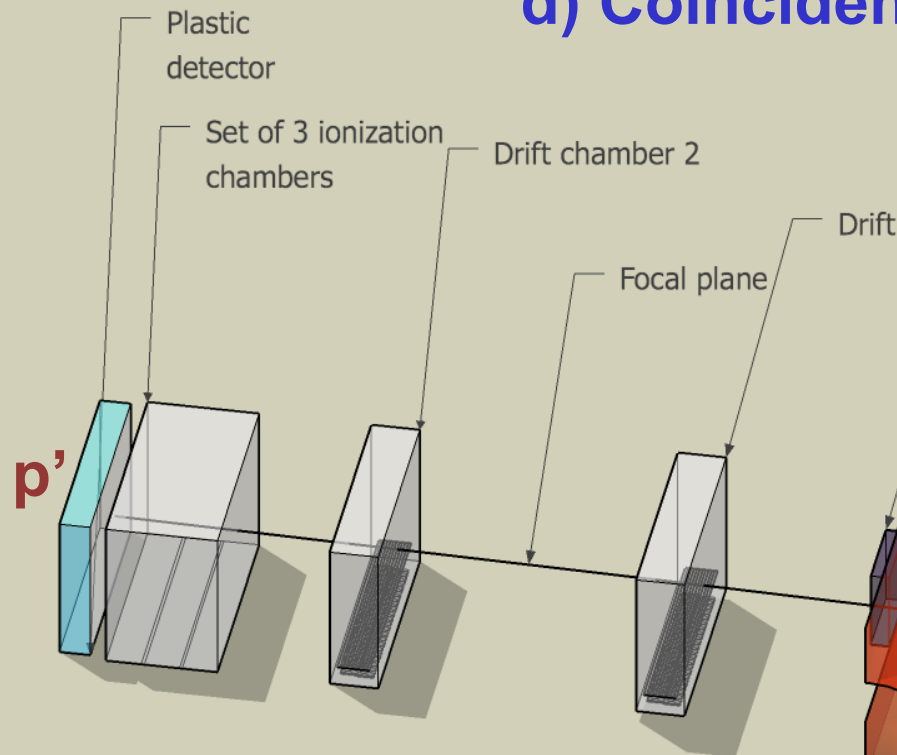


*T. Davinson *et al.* NIMA 454(2000) 8/13

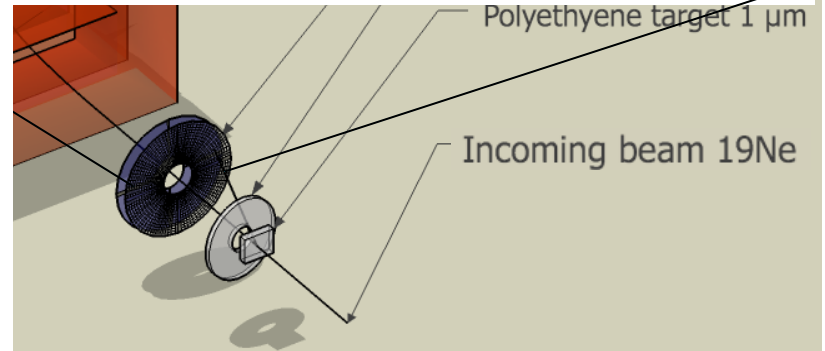
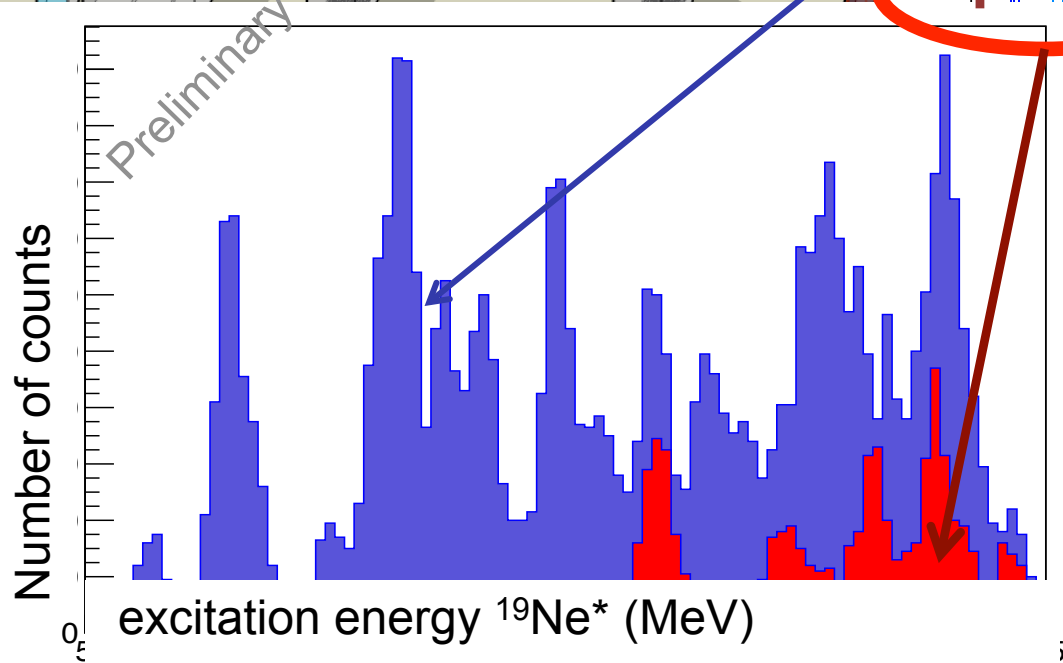
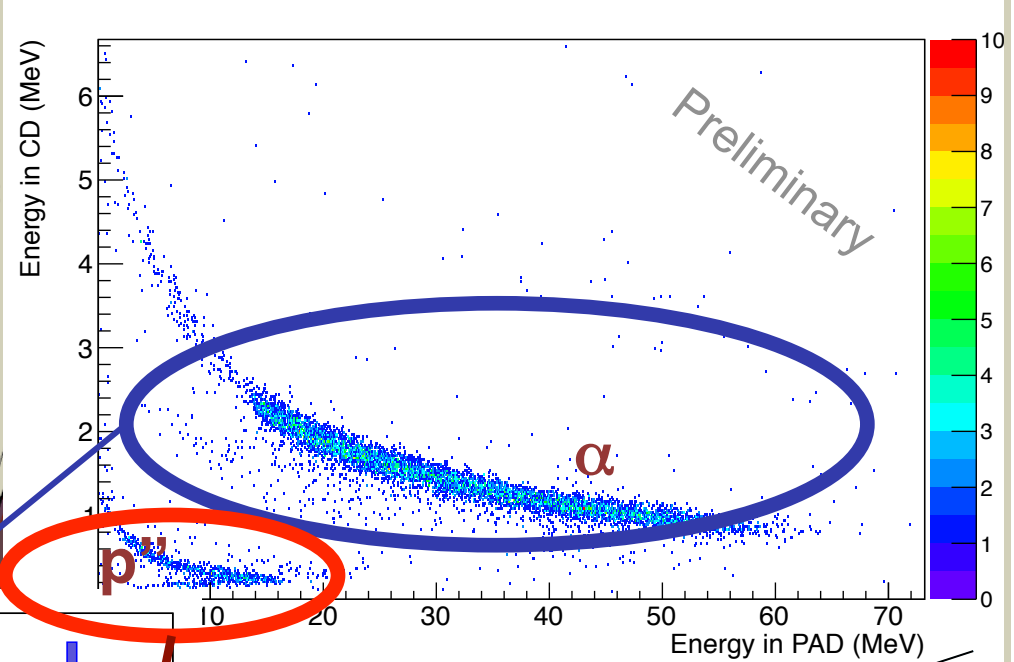
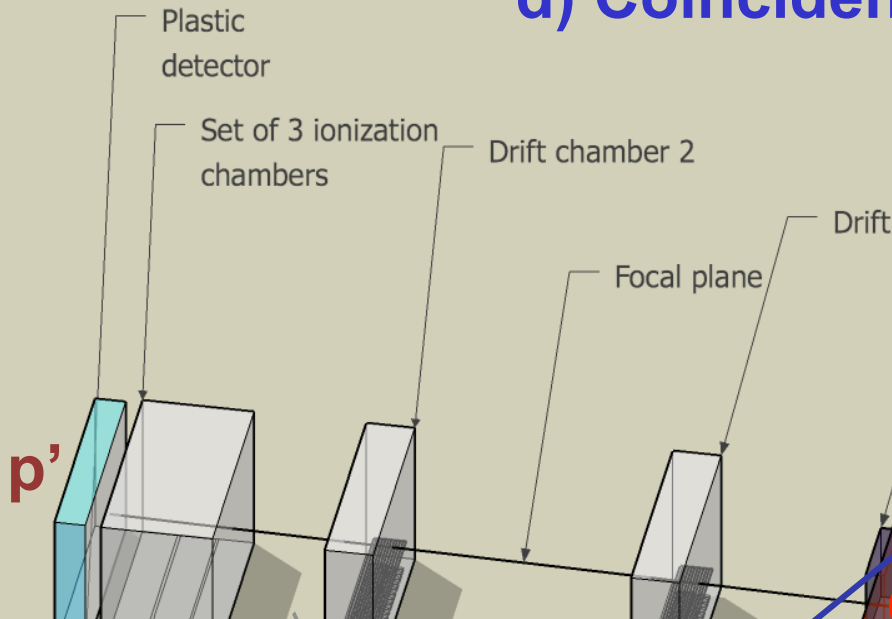
d) Coincidence with CD-Pad



d) Coincidence with CD-Pad



d) Coincidence with CD-Pad



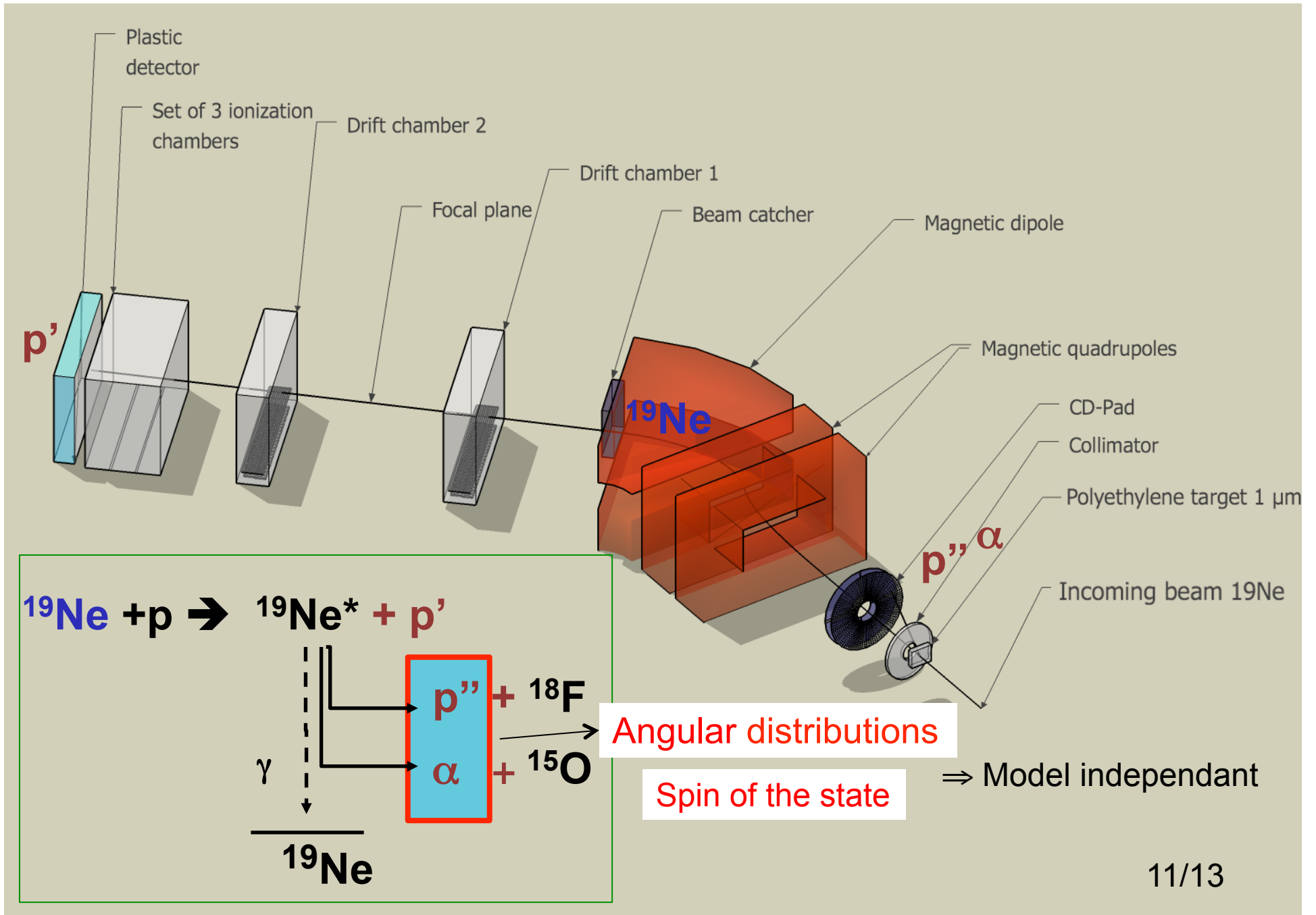
!!! Extract Γ values !!!

D.J. Mountford *et al.* PRC (2012)

1) Astrophysical context

2) The experiment

3) Angular distribution



1) Astrophysical context

2) The experiment

3) Angular distribution

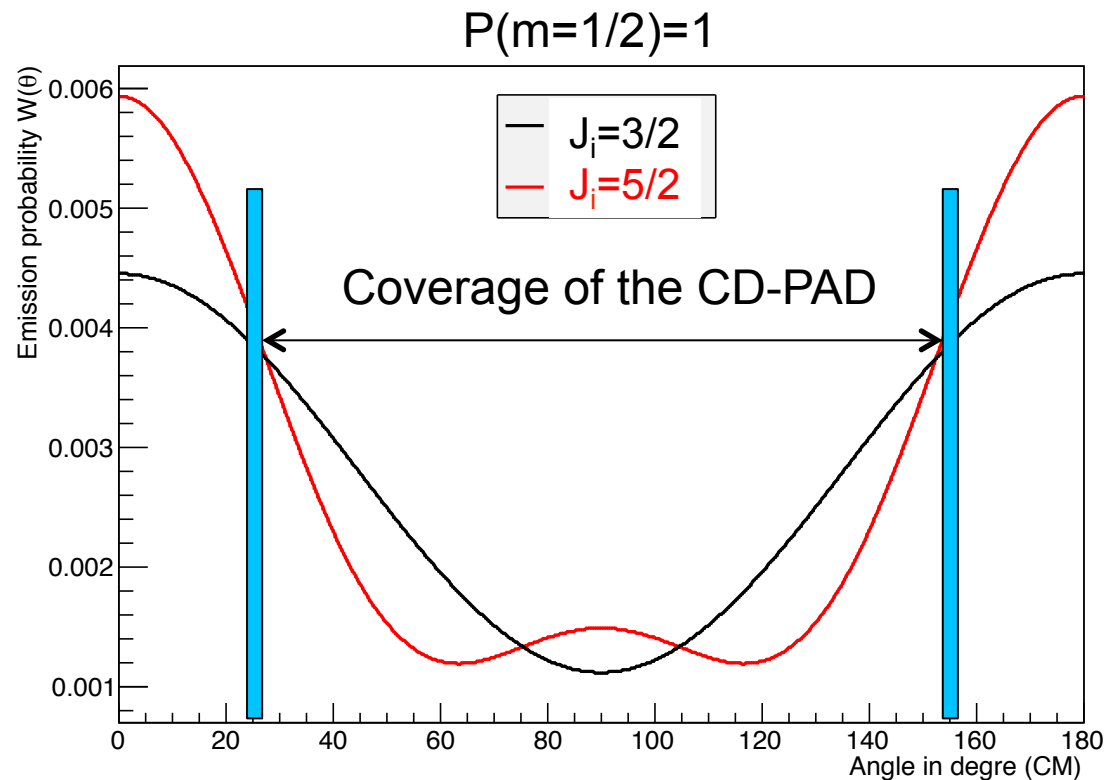
J.G. Pronko and R.A. Lindgren
Nuclear Instruments and Methods (1972)



$$W(\theta) = \sum_{mK} P(m) A(JI' smK) Q_K P_K(\cos\theta) \rightarrow \text{with } K = \min(l+l', 2J_i) \rightarrow \text{Spin}$$

where

$$A(JI' smK) = (-1)^{|s-m|} \hat{U} \hat{J}^2 (U' 00 | K0) \times (JJ m - m | K0) W(UI' J; sK)$$



Conclusion

Experiment

Online analysis : experiment is successful! ✓
(range of interest covered with good statistics...)

The resolution is better with this new inelastic scattering method. ✓

Access to spin and widths in a model independent way ✓

About the fine analysis

Study of angular correlation theory ✓

Starting of the analysis of the CD-PAD data

Thank you for your attention
A big thank to the collaboration !!!

GANIL (France) : F. Boulay, B. Bastin, F. De Oliveira, A. Lemasson, M. Rejmund, C. Schmitt, B. Jacquot, O. Kamalou, A.M. Sanchez Benitez, E. Traykov, C. Rodriguez, J. Grinyer, O. Sorlin, J.-C. Thomas and P. Delahaye.

University of Edinburgh (Scotland) : T. Davinson, V. Margerin, A. Estrade and P. J. Woods.

University of Santiago de Compostel (Spain) : D. Ramos.

University of York (England) : A. Laird.

IPN Orsay (France) : N. de Séréville.

University of Huelva (Spain) : G. Marquinez Duran and L. A. Acosta Sanchez.

LPC Caen (France) : L. Achouri.

Vinca Institute (Serbia) : P. Ujic.

Rez (Czech Republic) : J. Mrazek.

IFIN/HH (Romania) : F. Negoita, F. Rotaru, M. Stanoiu and C. Borcea .

Niewodniczanski Institute of Nuclear Physics (Poland) : M. Ciemala.