

11<sup>th</sup> Russbach School on Nuclear Astrophysics

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Chemo-Dynamical  
evolution of dwarf  
spheroidal galaxies

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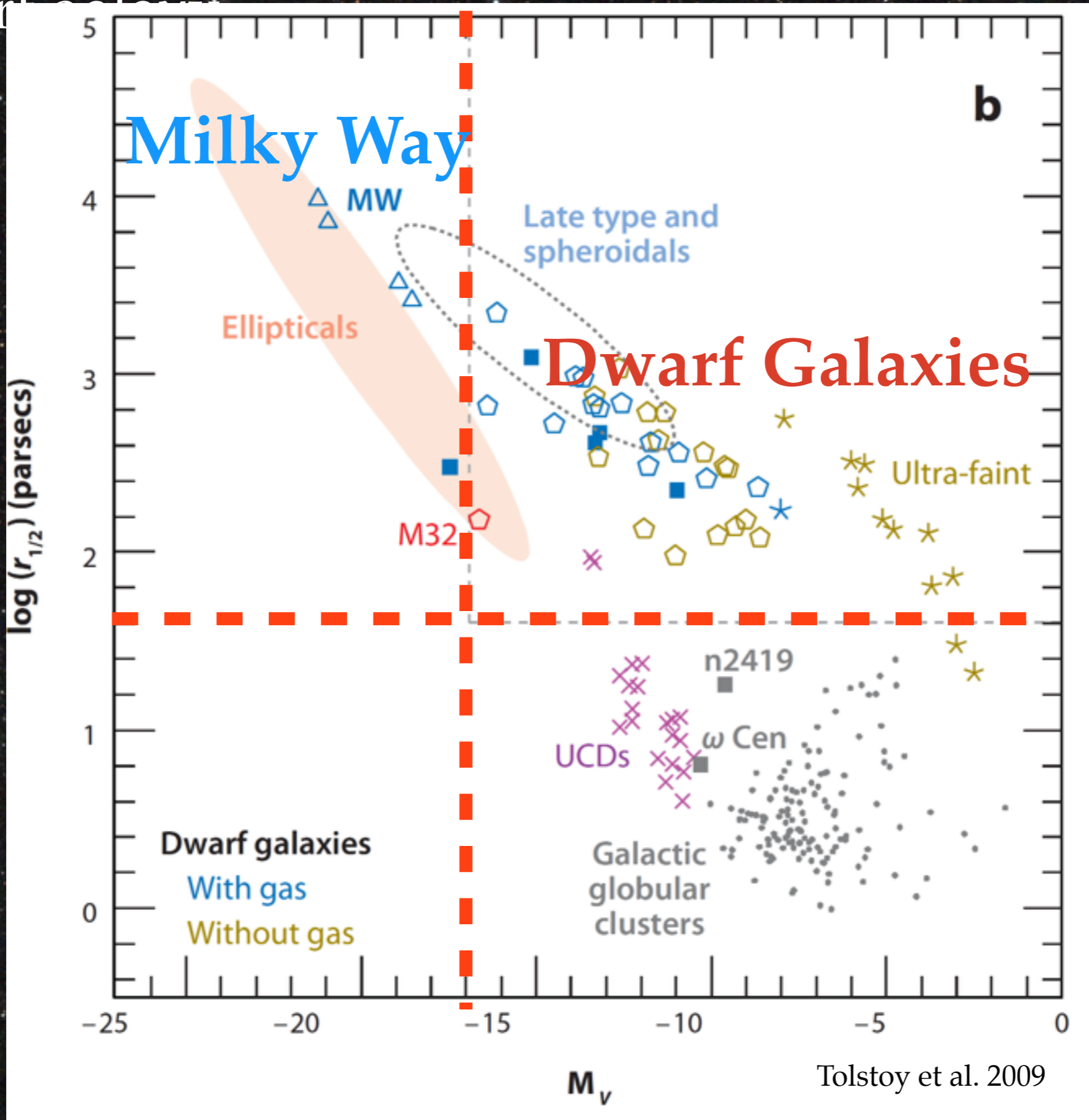
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Grant J. Mathews (U. Notre Dome),

and Toshitaka Kajino (NAOJ, UTokyo)

# Dwarf Galaxies

Leo II dwarf galaxy



# Galaxy Formation

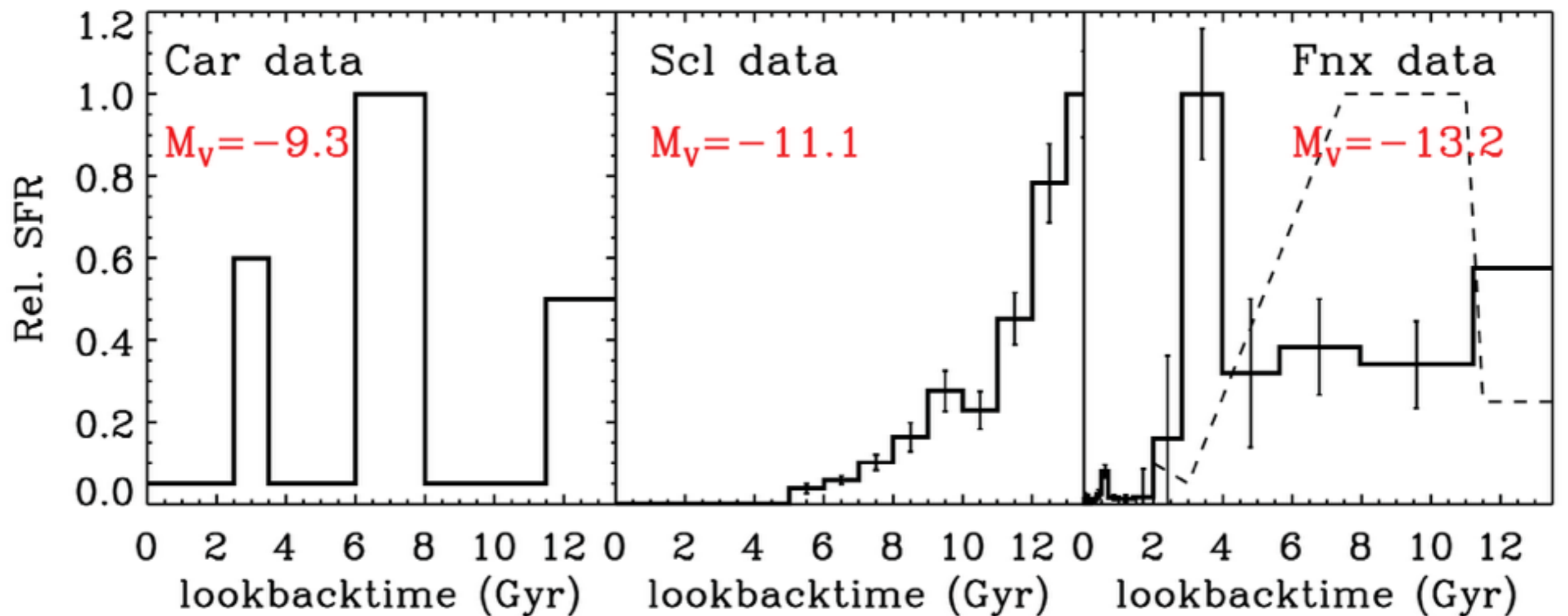
Dwarf Galaxies might be building  
blocks of the Milky Way

We can observe each star in nearby  
dwarf galaxies to estimate star  
formation histories and metallicity.

Dwarf galaxies are useful tool to study galaxy  
formation and evolution

# Observation

## Star Formation Histories of dwarf galaxies

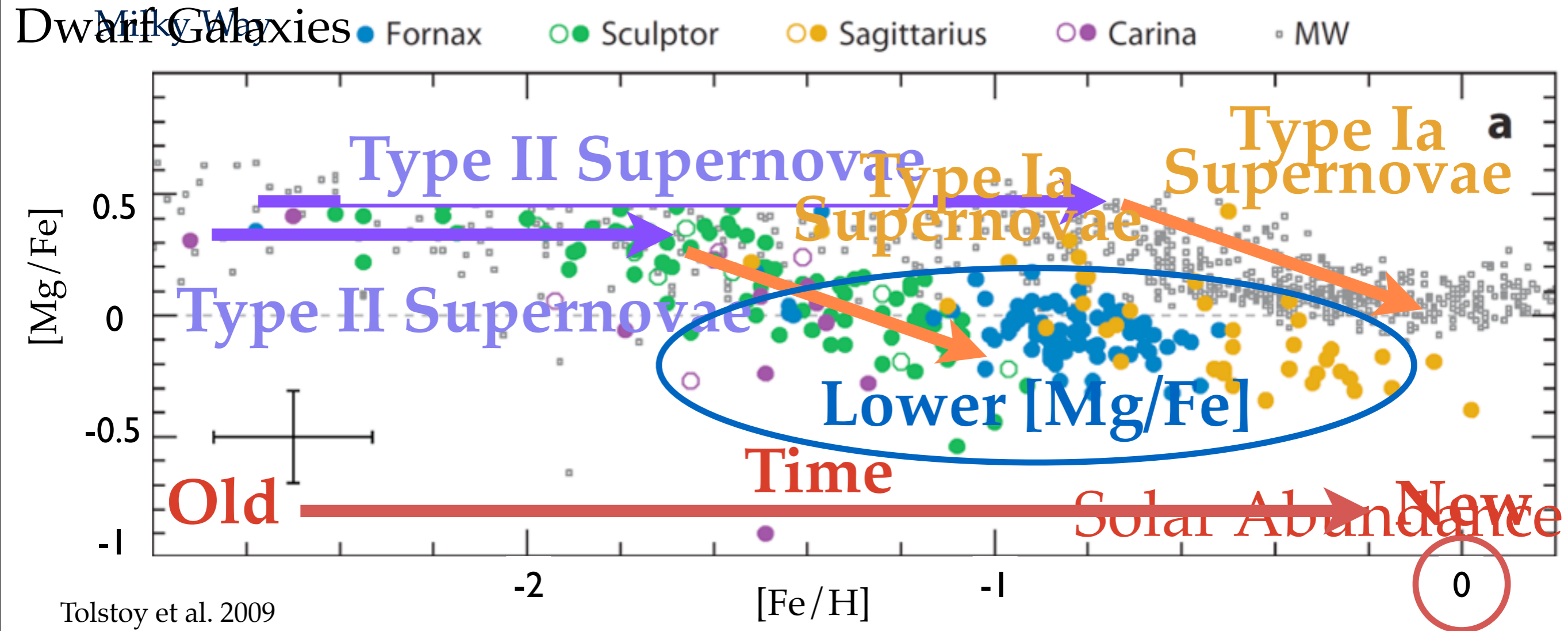


Starckenburg et al. 2013

Each dwarf galaxies might experience different evolution.

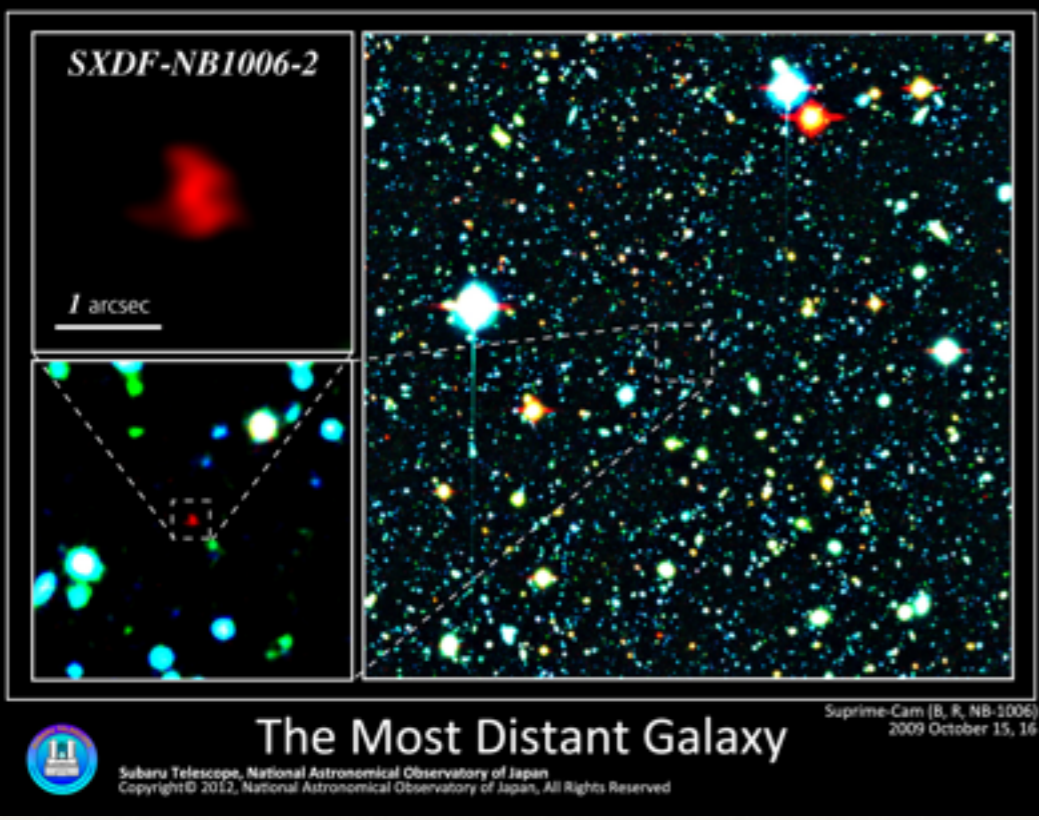
Star formation histories affect the chemical abundance pattern of galaxies

# Elemental Abundance Pattern of dwarf galaxies

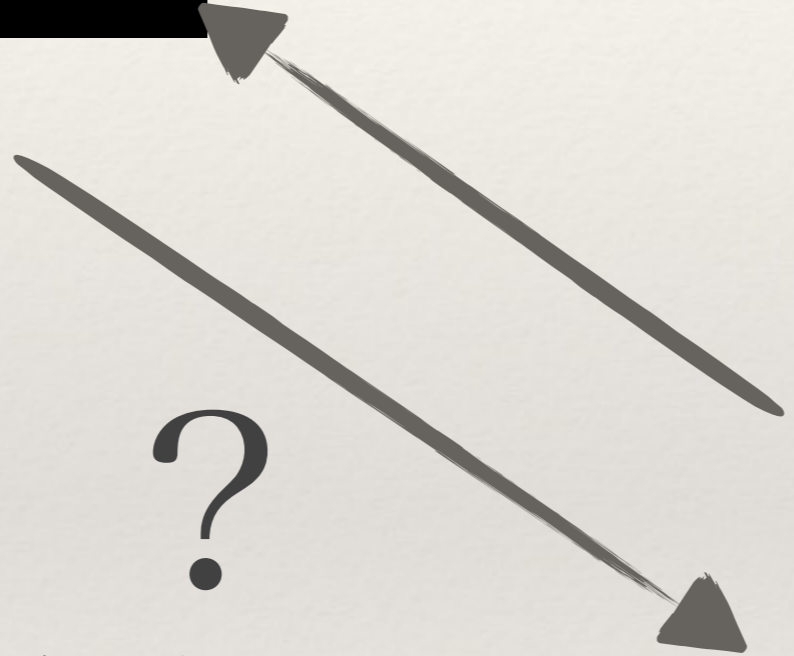


$$[\text{Fe}/\text{H}] = \log_{10} \left( \frac{N_{\text{Fe}}}{N_{\text{H}}} \right)_{\text{star}} - \log_{10} \left( \frac{N_{\text{Fe}}}{N_{\text{H}}} \right)_{\text{sun}}$$

Systematically different chemical abundance pattern from Milky Way



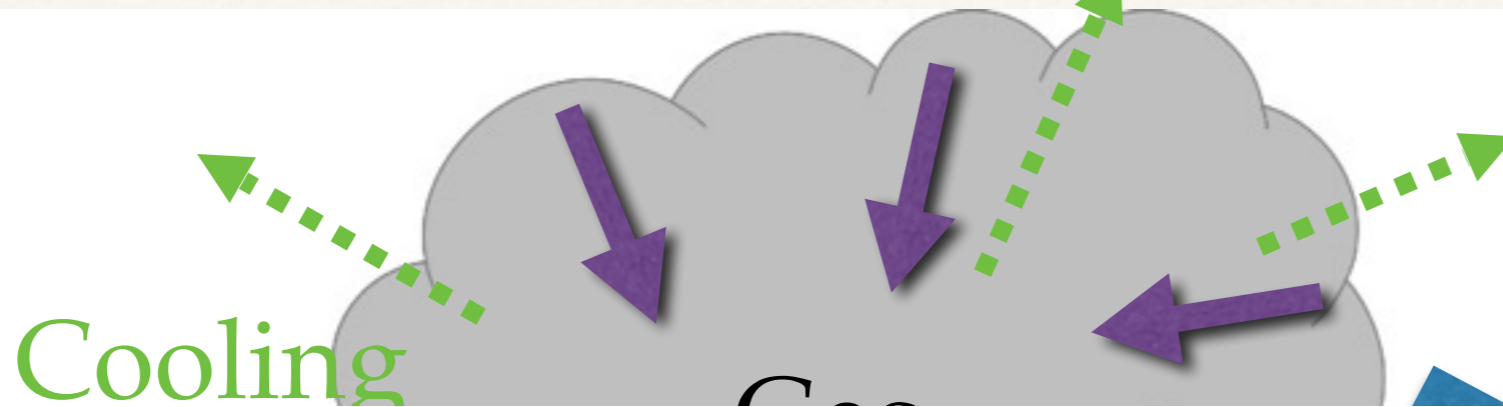
Present dwarf galaxies and building blocks of the Milky Way are different?



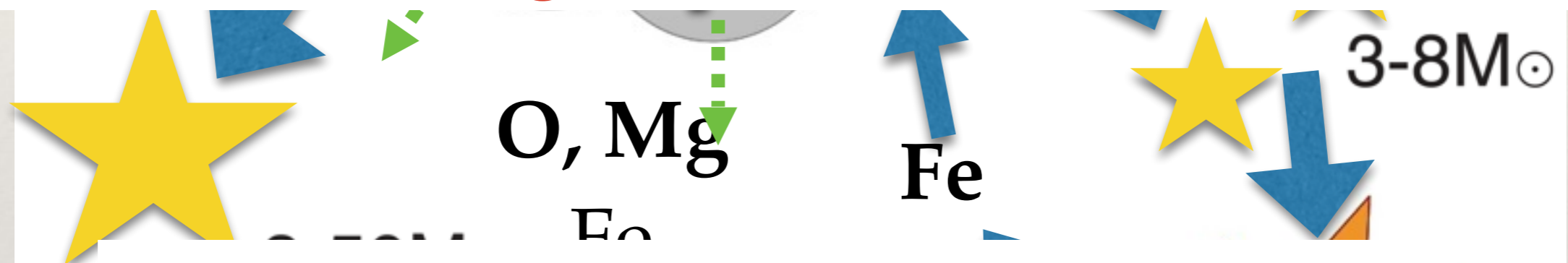
We need detailed simulation of dwarf galaxy evolution



# Physical Processes inside a galaxy



We need to simulate dynamical and chemical evolution of galaxies at the same time!



Chemical and Dynamical Evolution

= Chemodynamical evolution

Massive Stars

Type II Supernovae

Type Ia  
Supernova

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# Our Ultimate Goal

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to reveal the chemo-dynamical evolution of dwarf galaxies in order to construct a comprehensive picture of the formation and evolution of the Milky Way in terms of the origin and evolution of atomic elements!



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# Today's goal

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- to construct a “**chemo-dynamical evolution code**”
- to find important physics on **star formation histories**

## Next Goal

- to implement “supernova nucleosynthesis yields” which we are calculating in our Tokyo Group into the present chemo-dynamical code
- to study the “chemical evolution” and “dynamical evolution” of dwarf spheroidal galaxies simultaneously as building blocks of understanding the Milky Way



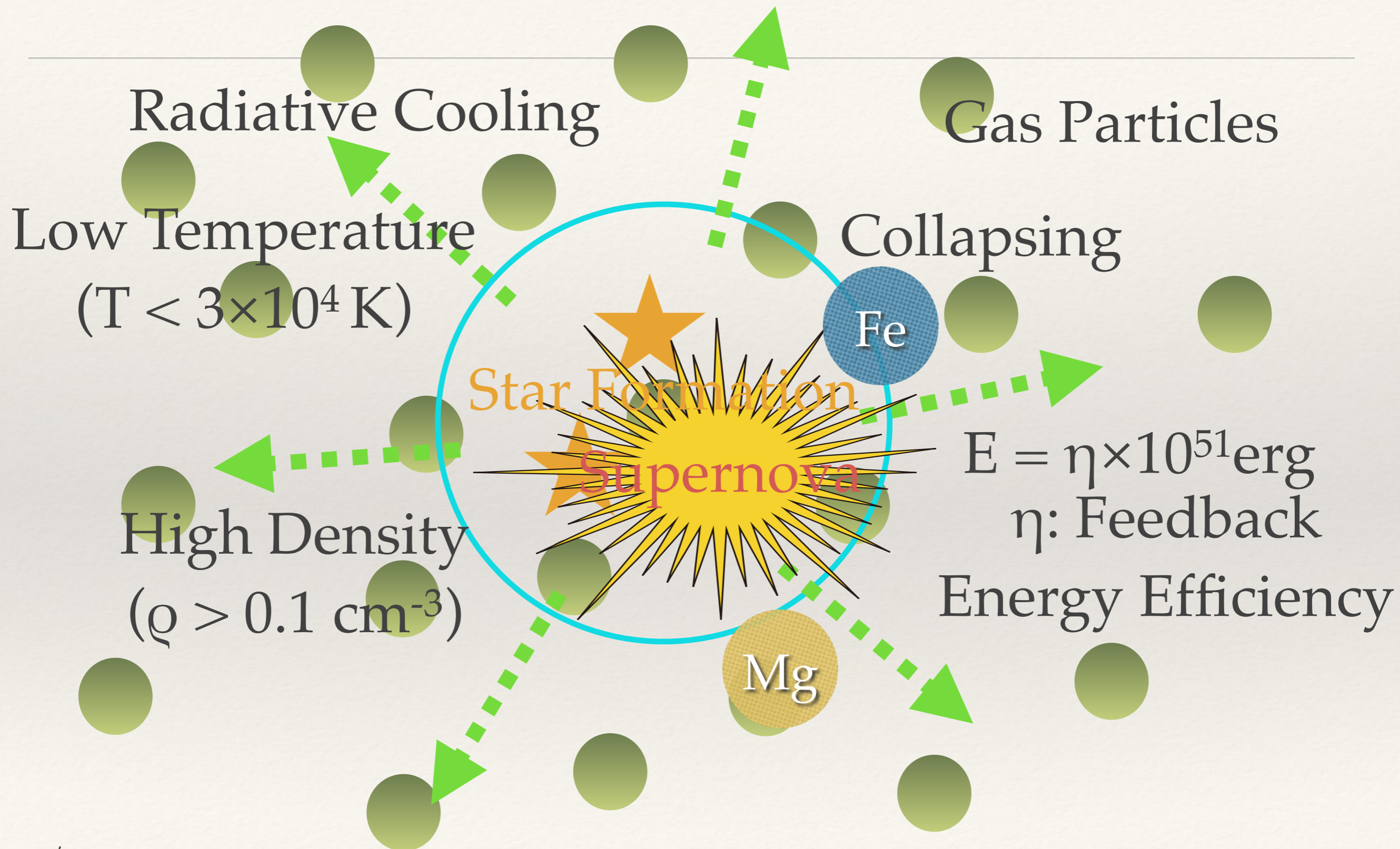
# Chemo-Dynamical Evolution Code

Dark Matter (Gravity, Tree method)



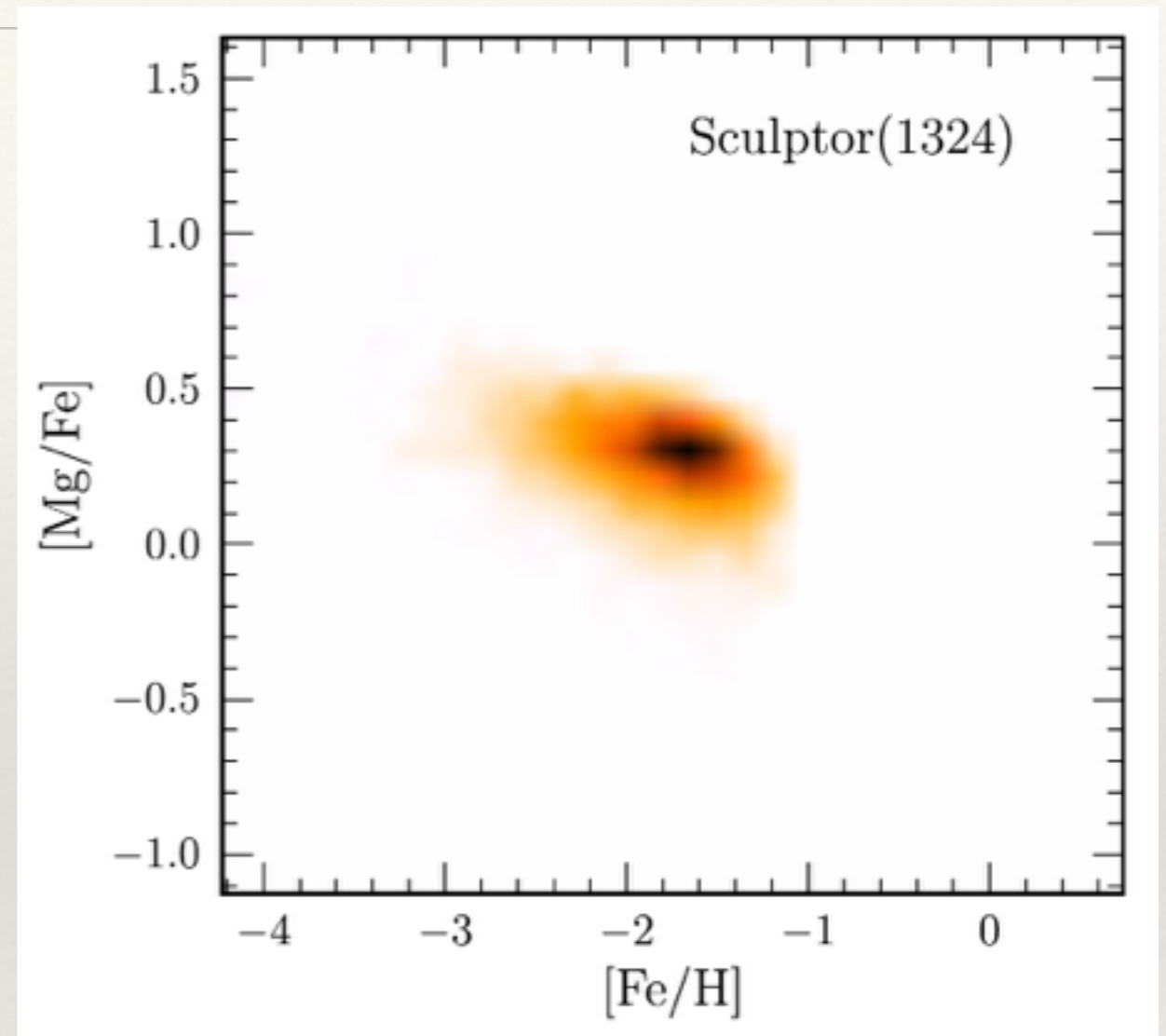
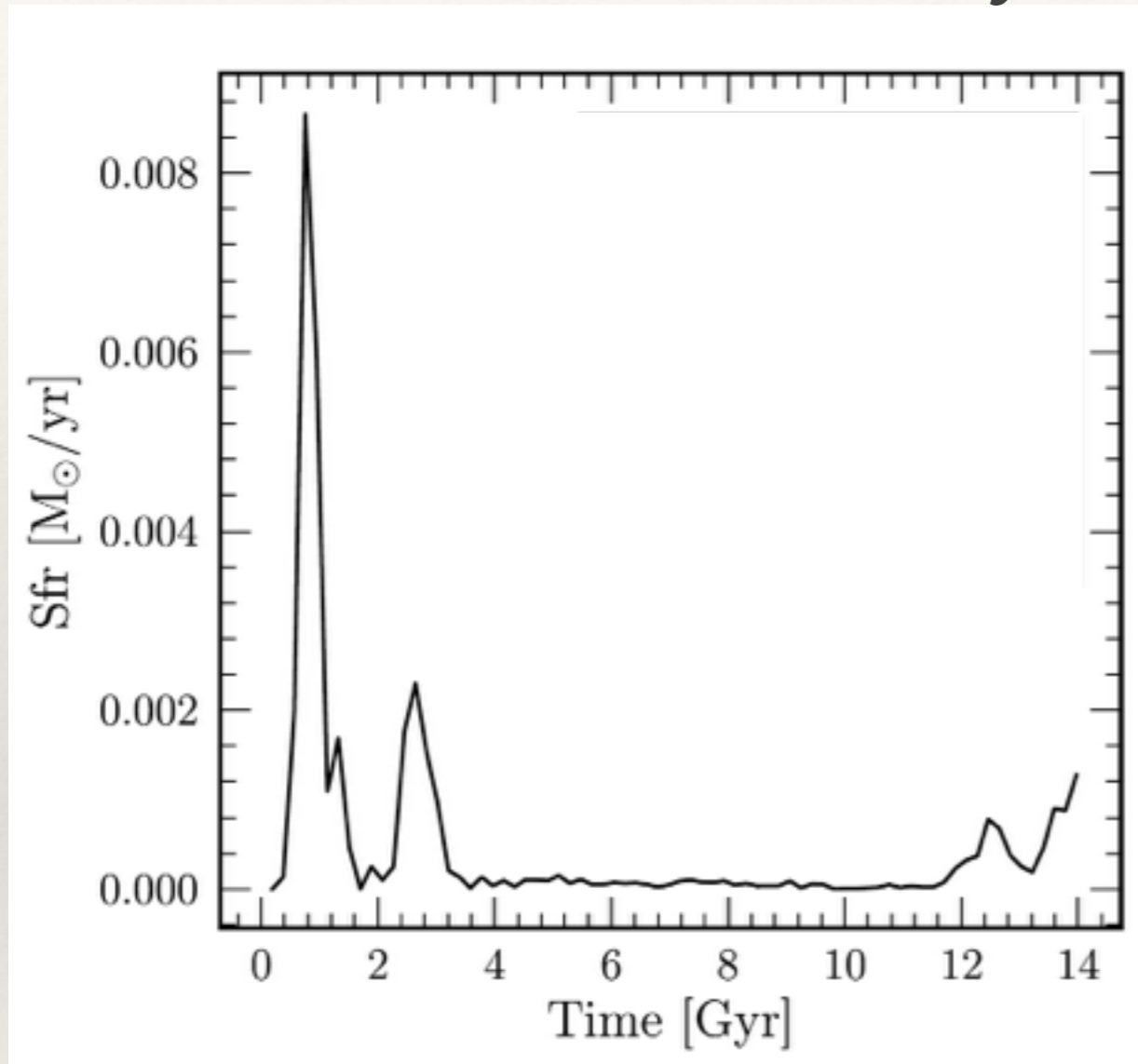
Gas (Hydrodynamics,  
SPH method, ASURA  
(Saitoh & Makino, 2013))

# Star Formation Law and Supernova Feedback



# Past Studies (Revaz & Jablonka 2012)

Star Formation History    Chemical Abundance Pattern



Successfully reproduce star formation histories and  
chemical abundance patterns

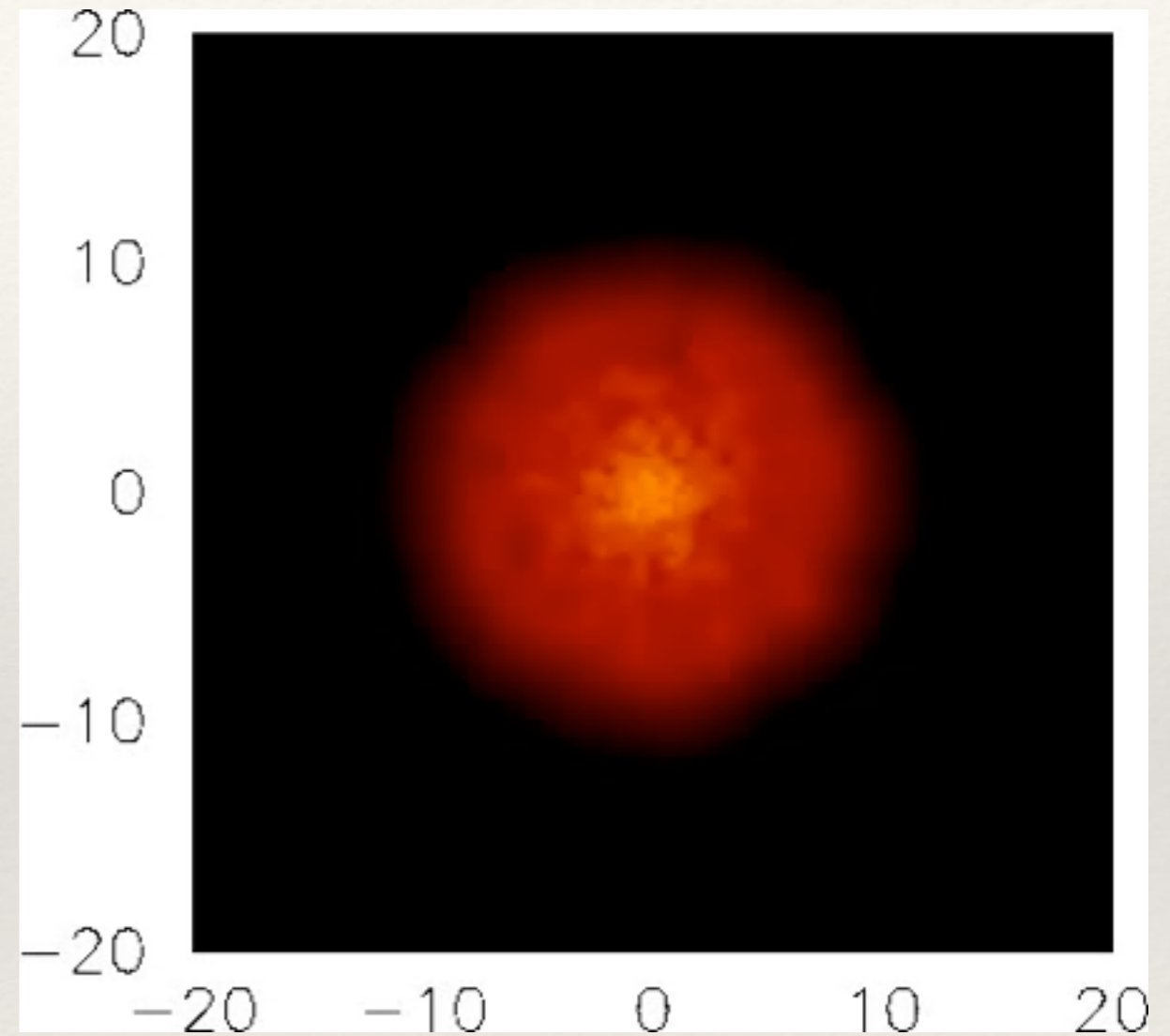
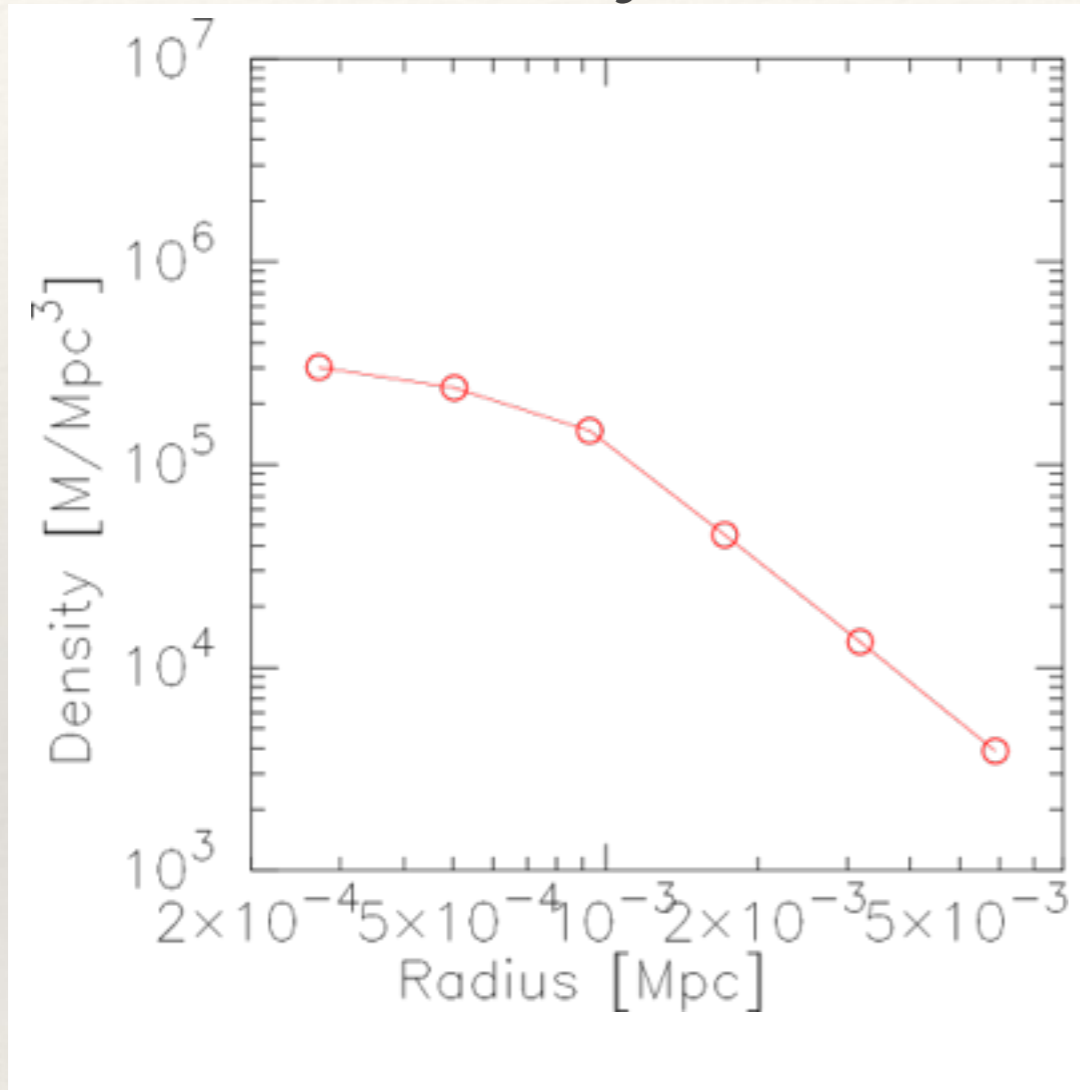
But... they reduce the energy of the supernova  
explosion ( $10^{49}$  erg)

# Initial Condition

Density Profile: pseudo-isothermal profile

Radial Density Profile

Density Distribution

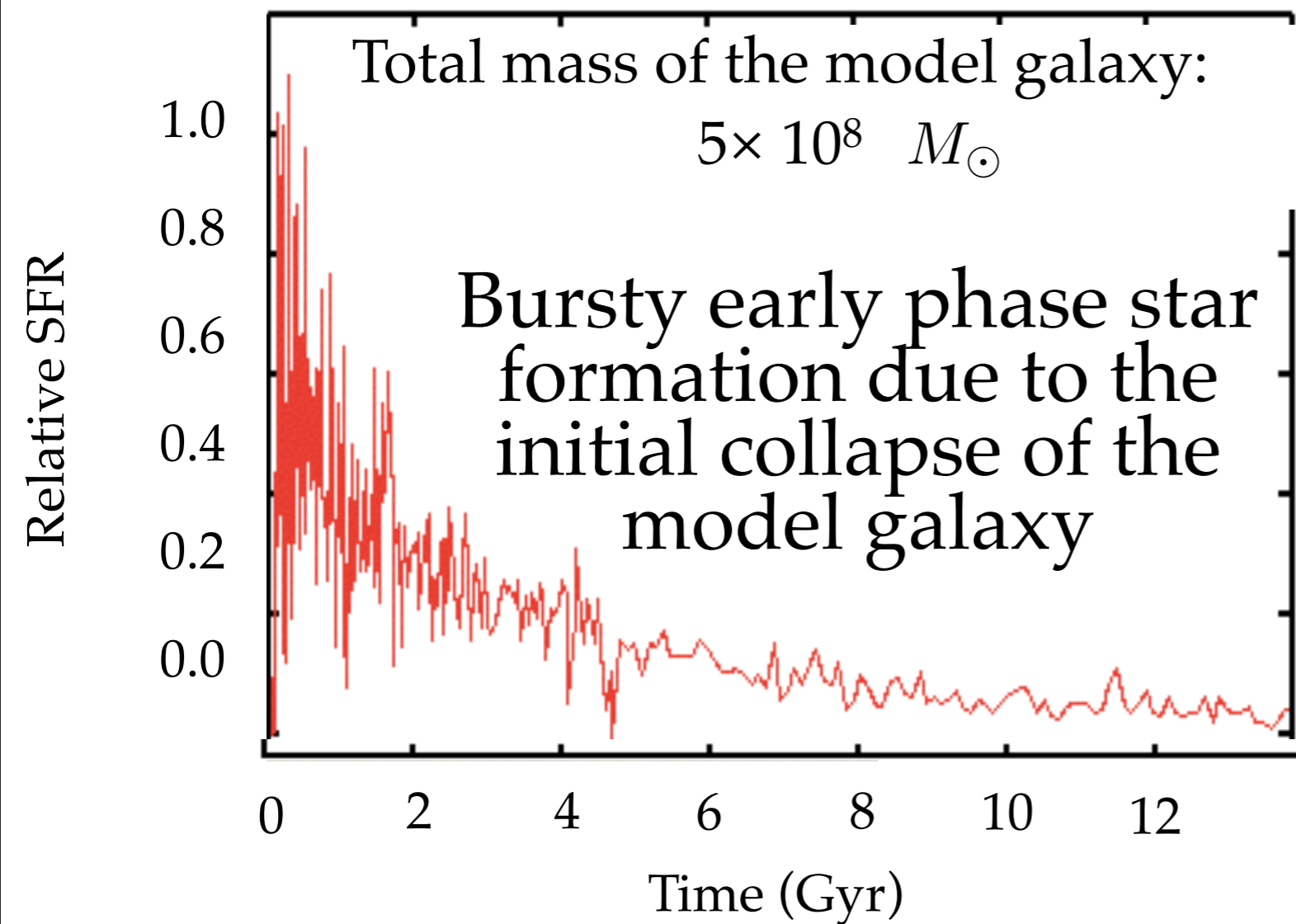


Total Number of Particles:  $2^{16}$

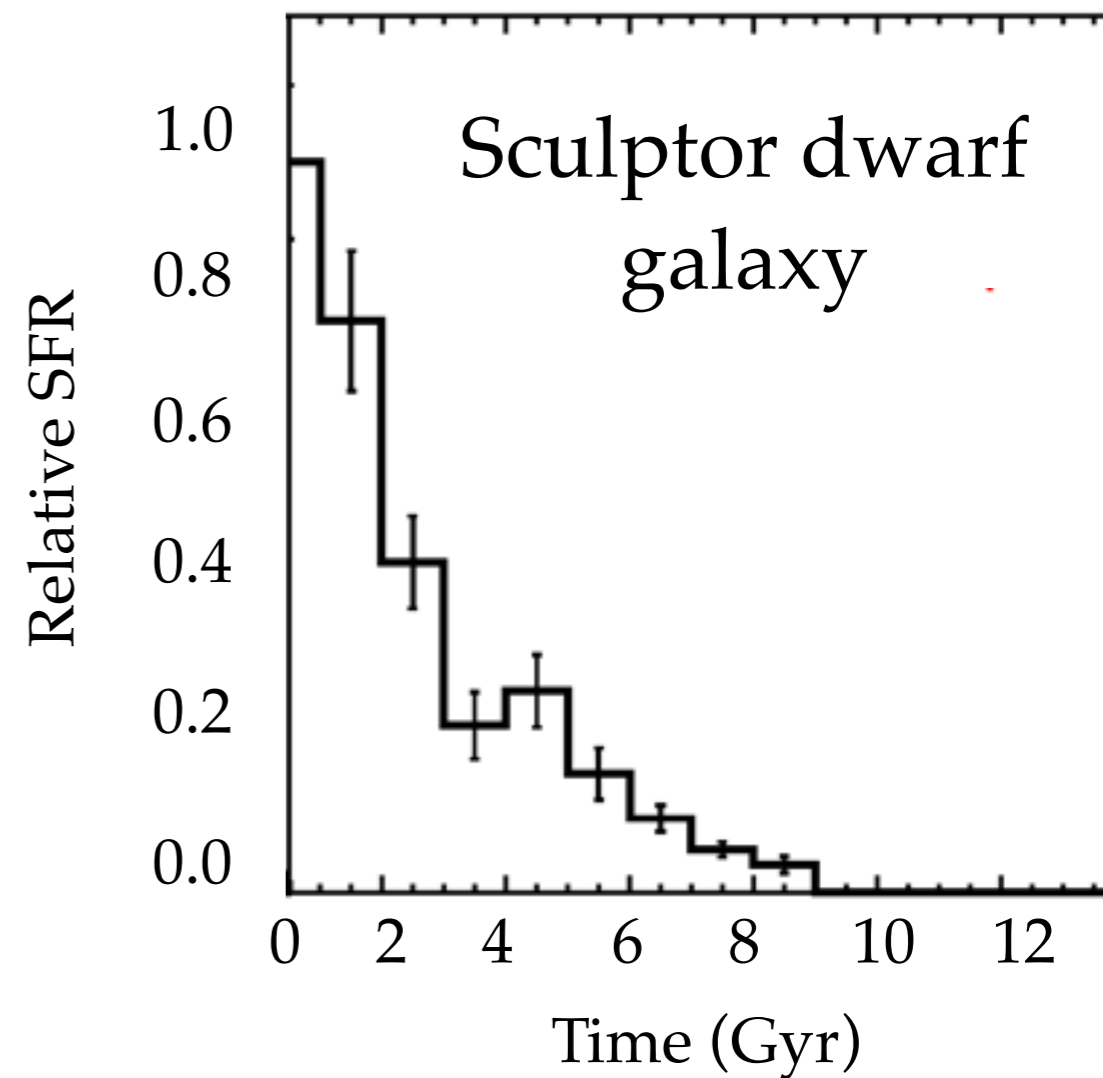
Mass of One Gas Particle:  $10^3 M_{\odot}$

# Star Formation Histories

## Simulation

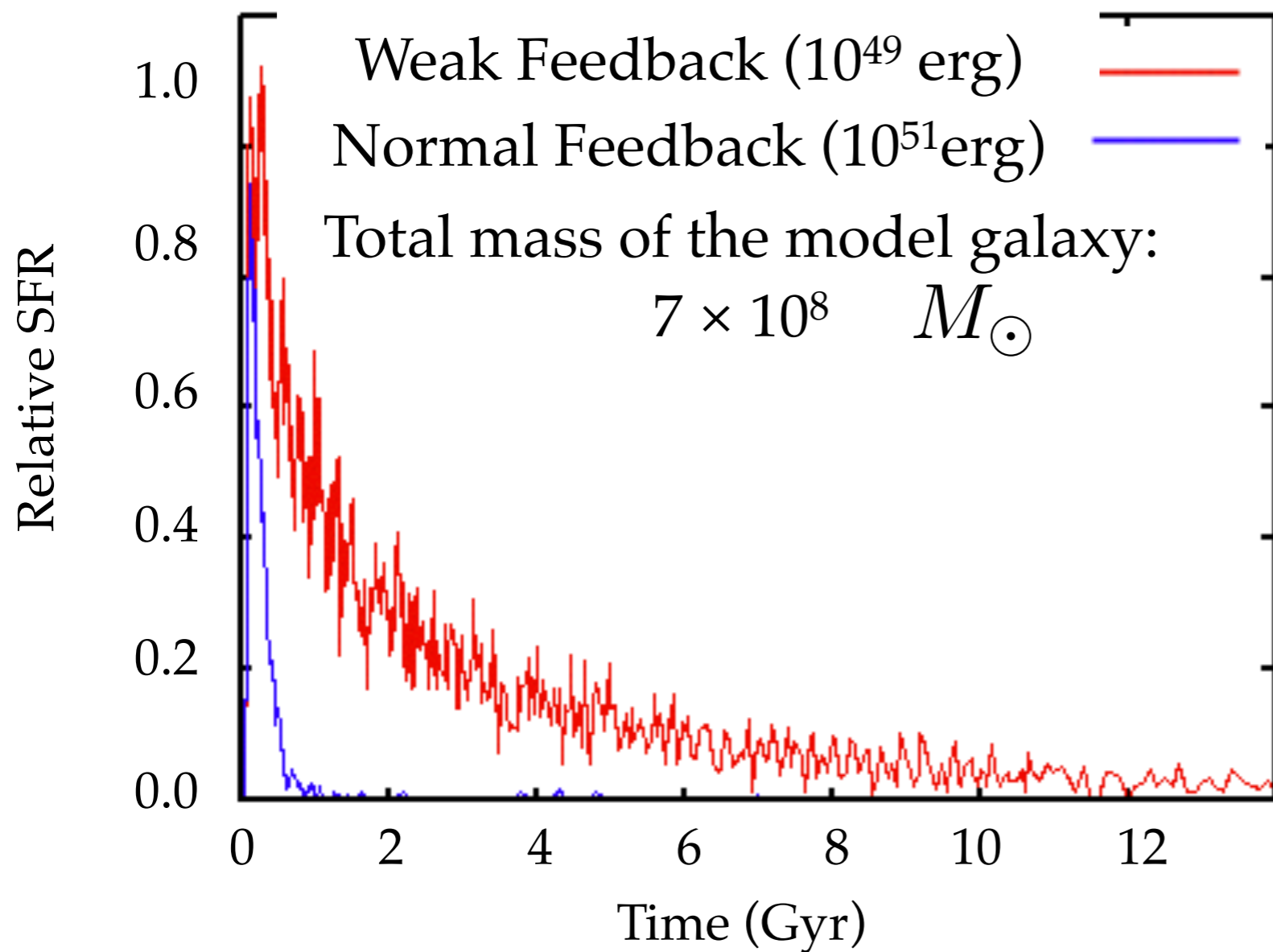


## Observation



Sculptor dwarf galaxy seems to evolve isolated

# Importance of supernova feedback



- ❖ Supernova feedback strongly affects the star formation rate
- ❖ Strong feedback blew away the gas inside the galaxy
- ❖ Chemical abundance pattern may be affected by supernova feedback

For near future

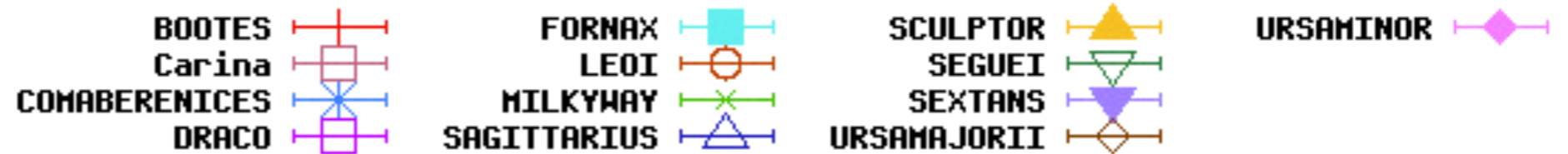
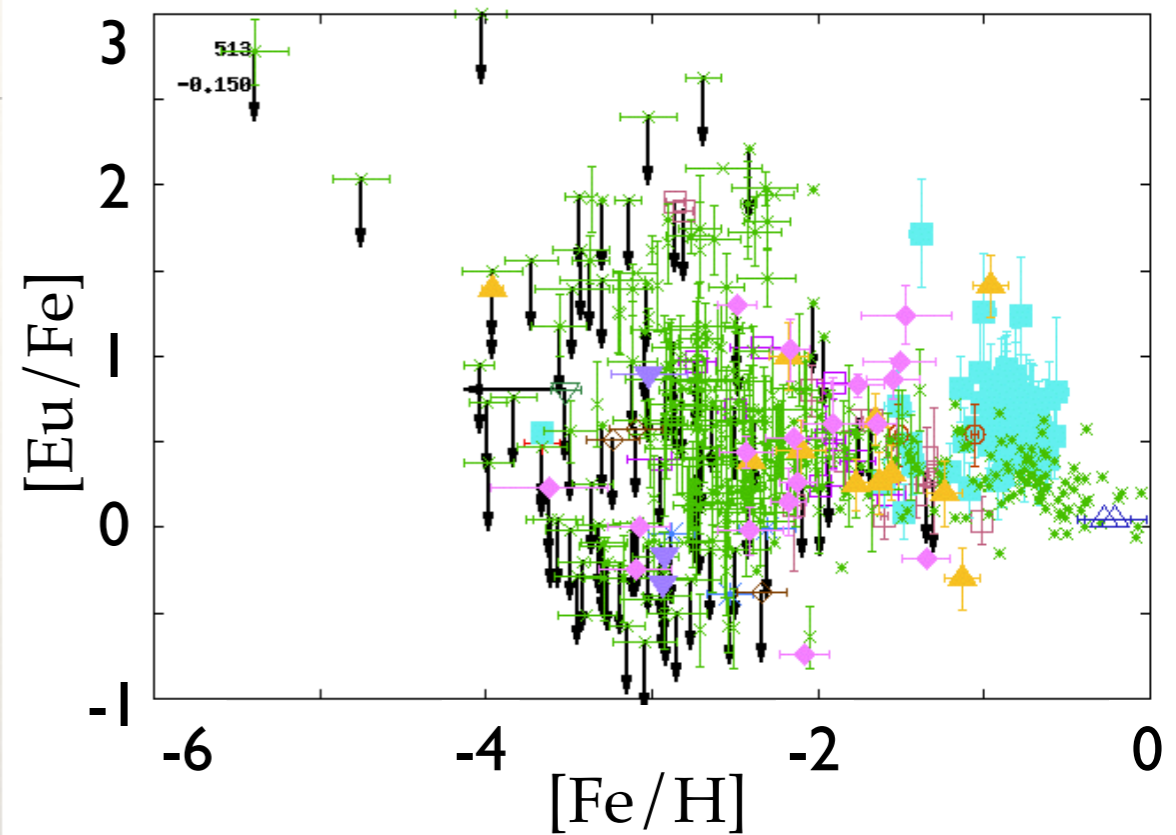
We need to check more massive galaxy and effects of merger with feedback



# Next Prospects

Constraining the sites of r-process elements

Origin of r-process elements:  
Type II supernovae?  
Neutron Star Mergers?



SAGA database (Suda et al. 2008)

The answer might be in chemodynamical simulation of dwarf galaxies!

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# Summary

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- ❖ Dwarf spheroidal galaxies is useful tool to study galaxy formation and evolution.
- ❖ We constructed chemical and dynamical evolution code.
- ❖ Supernova feedback is an important process to derive star formation histories.
- ❖ We will include chemical feedback such as r-process elements to deeply study evolution of dwarf galaxies and origin of elements