



# Reaction Rates for Explosive Nuclear Synthesis

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# X-Ray Bursts – short intro

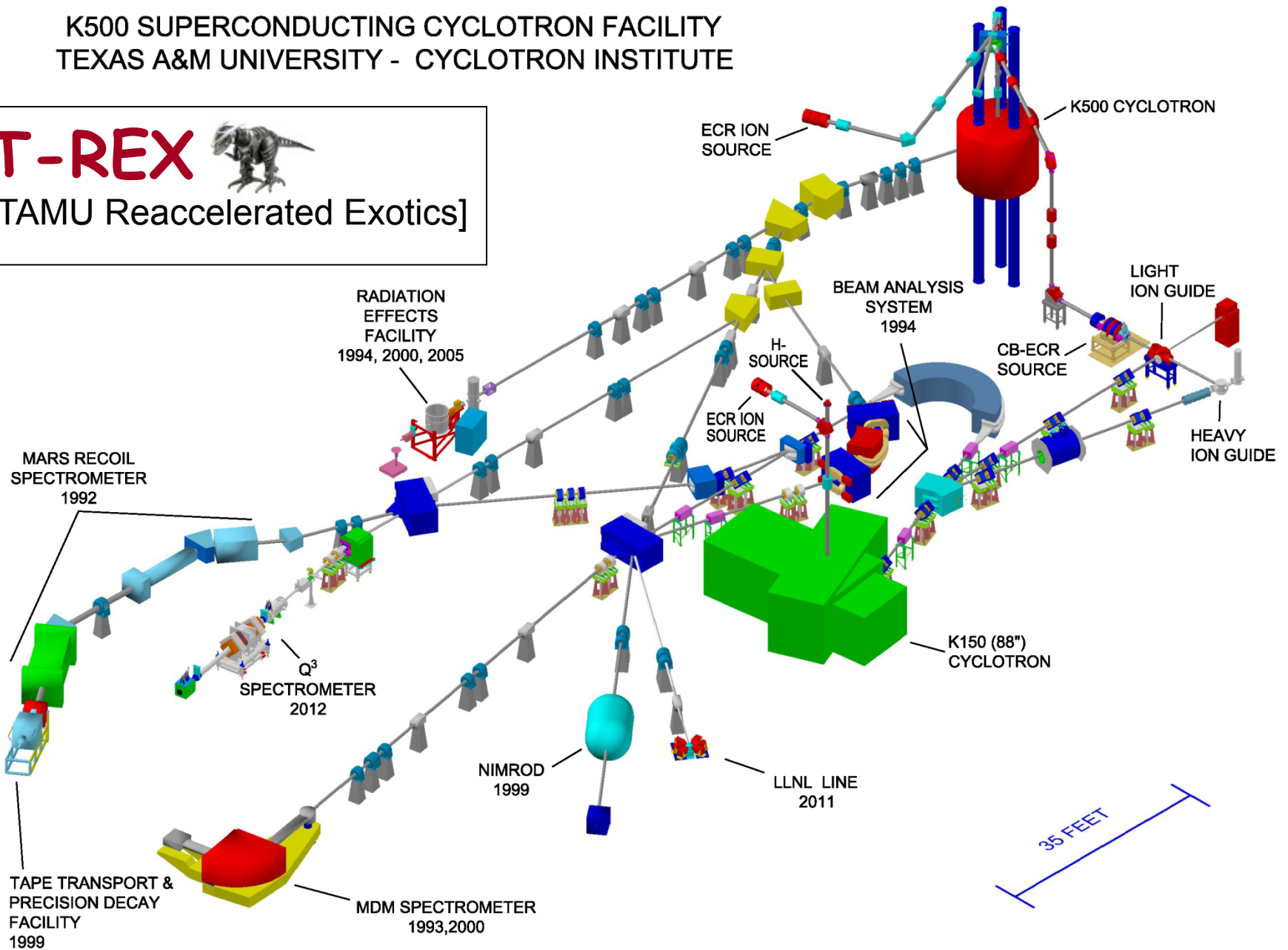
- Most frequent thermonuclear explosions in the universe
- Over 90 Galactic X-ray bursting sources detected to date
- Provide a unique window into the physics of neutron stars
- With the recent advances in observational astrophysics there is a large amount of data
- Problem:
  - Need a reliable nuclear physics database to interpret it
  - In particular, need good understanding of the rp-process

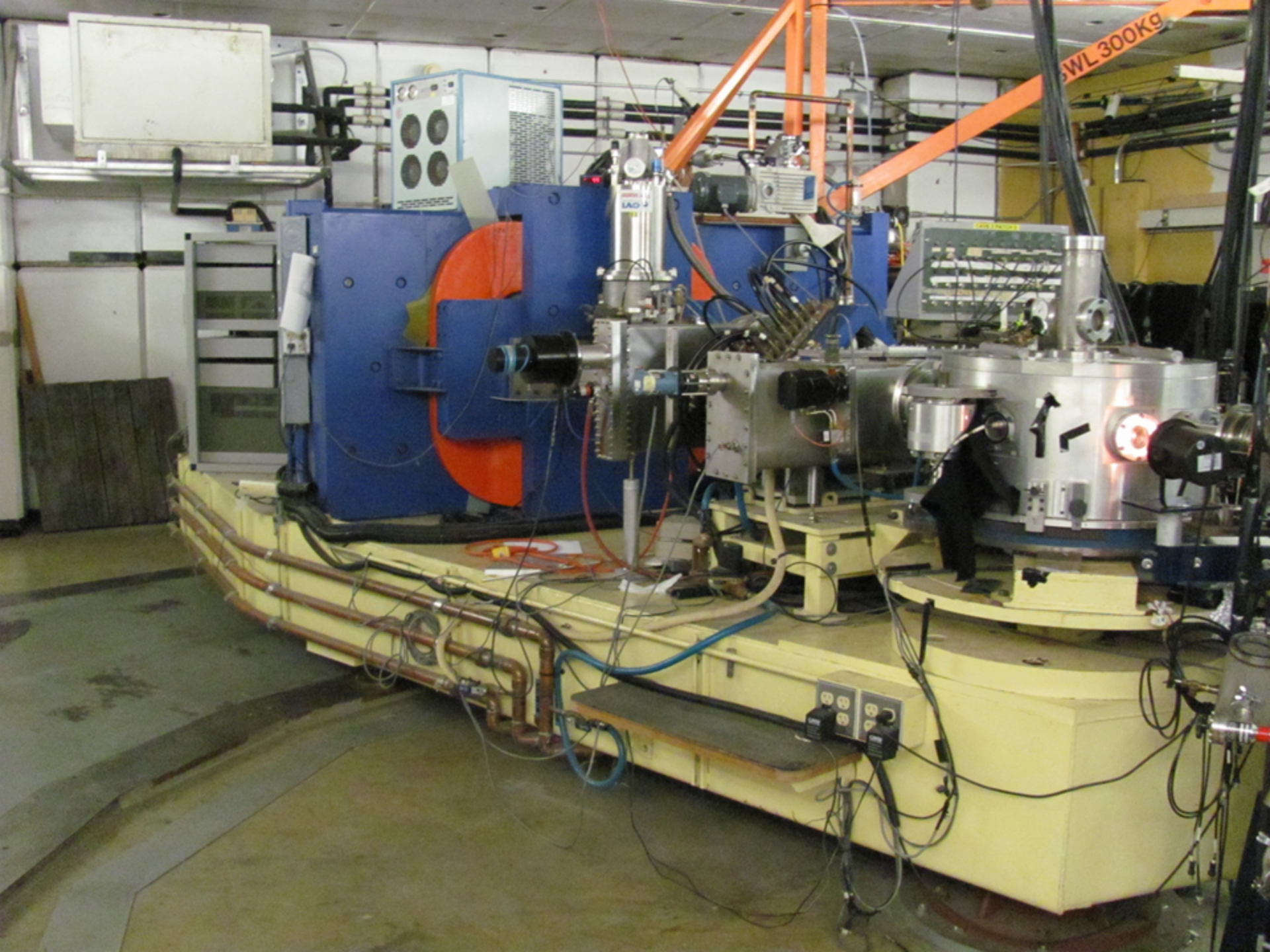
# X-Ray Bursts – short intro

- The critical nuclear data in the rp-process:
  - nuclear masses
  - $\beta$ -decay rates
  - nuclear reaction rates
- Most nuclei in rp-process are unstable
- Indirect methods have been used
- Large uncertainties
- Radioactive beams needed

# K500 SUPERCONDUCTING CYCLOTRON FACILITY TEXAS A&M UNIVERSITY - CYCLOTRON INSTITUTE

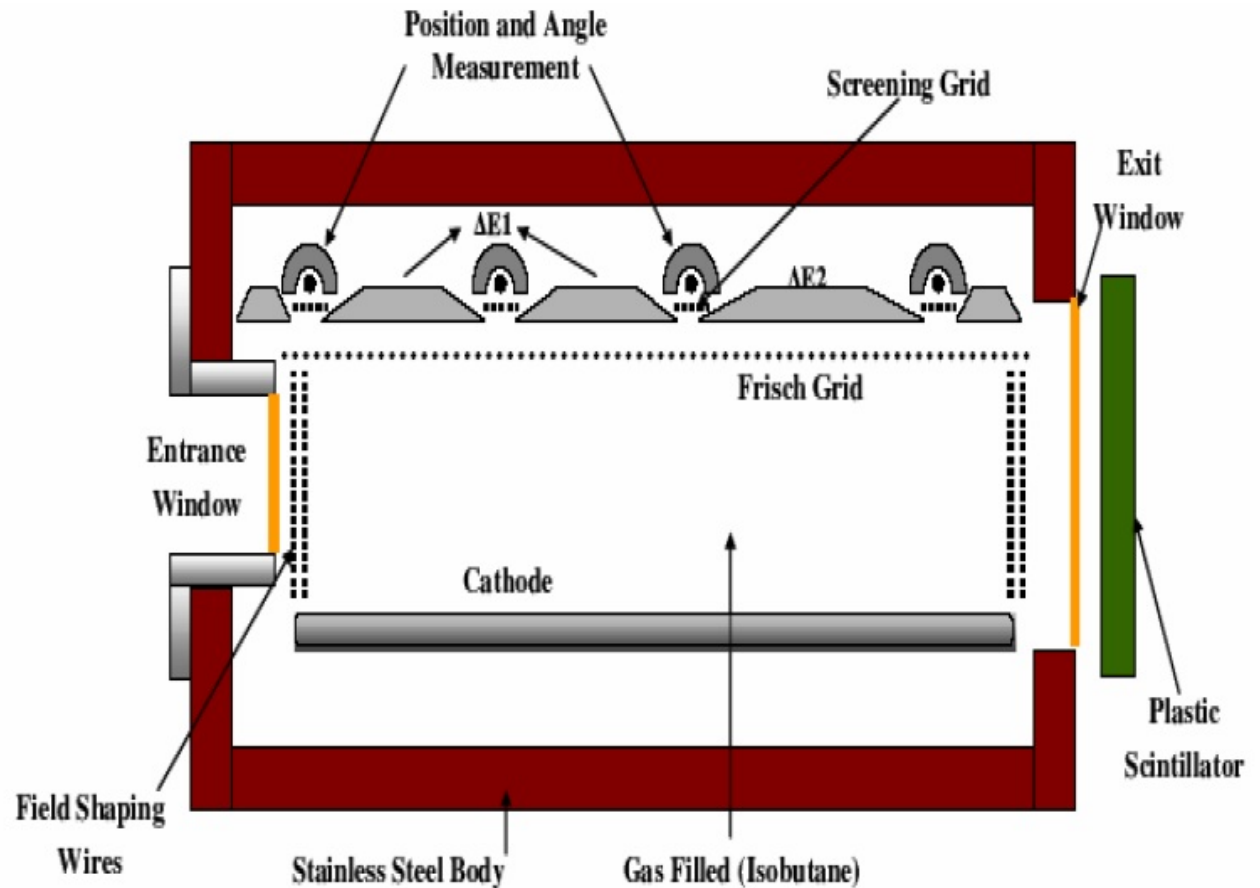
**T-REX**   
[TAMU Reaccelerated Exotics]





# MDM-Oxford detector

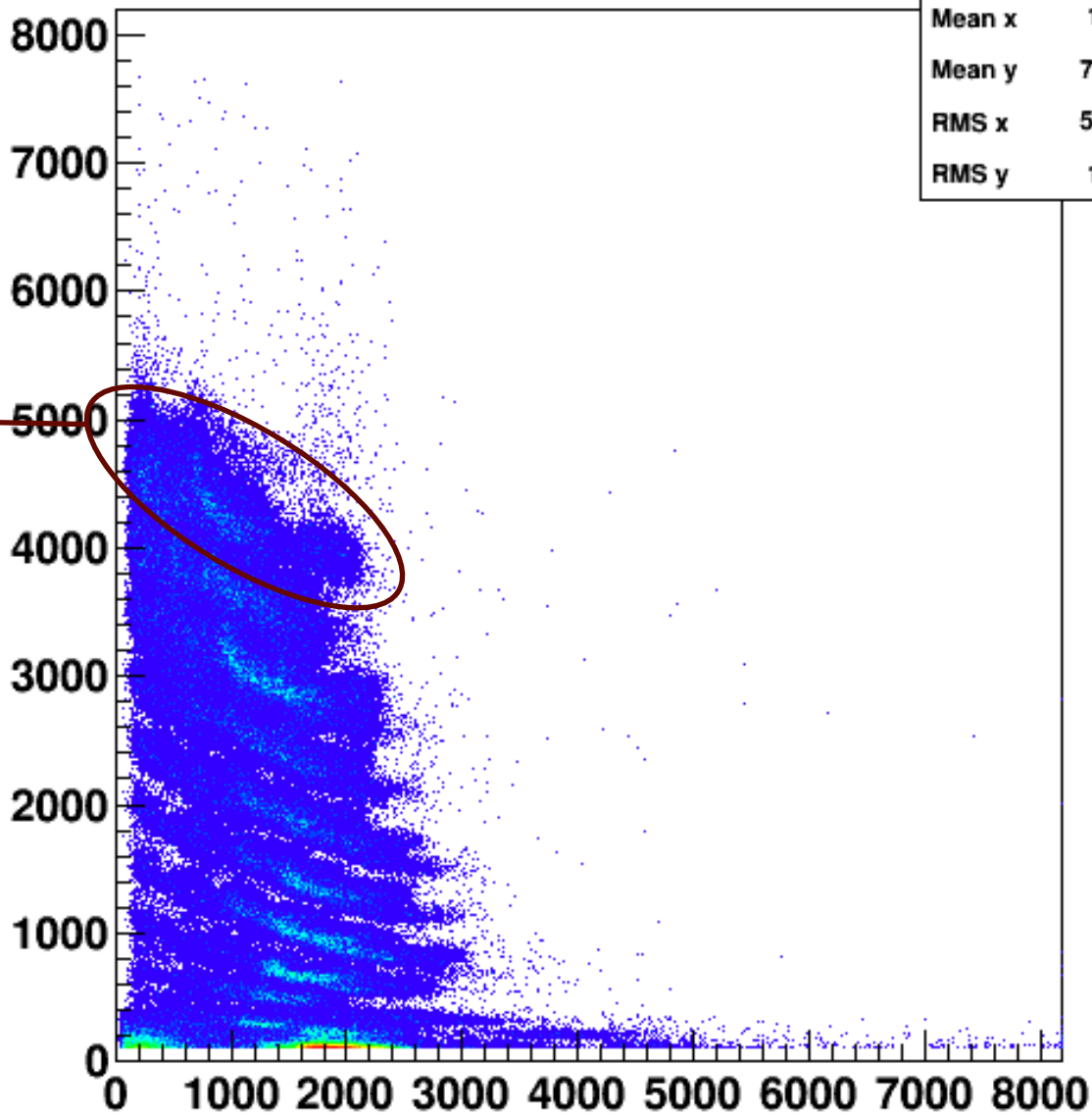
- Oxford detector
- Gridded ion chamber proportion
- Plastic scintillator PMTs for



DE1-PM

DE1-PM	
Entries	8943124
Mean x	1608
Mean y	770.5
RMS x	557.7
RMS y	1248

Mg isotopes





Micromegas

Micro Mesh Gaseous Structure, Y. Giomataris, Ph. Rebourgeard, J-P Robert and G. Charpak, NIM A376, 1996, p29 (CEA-biospace patent)

MICROMEGAS

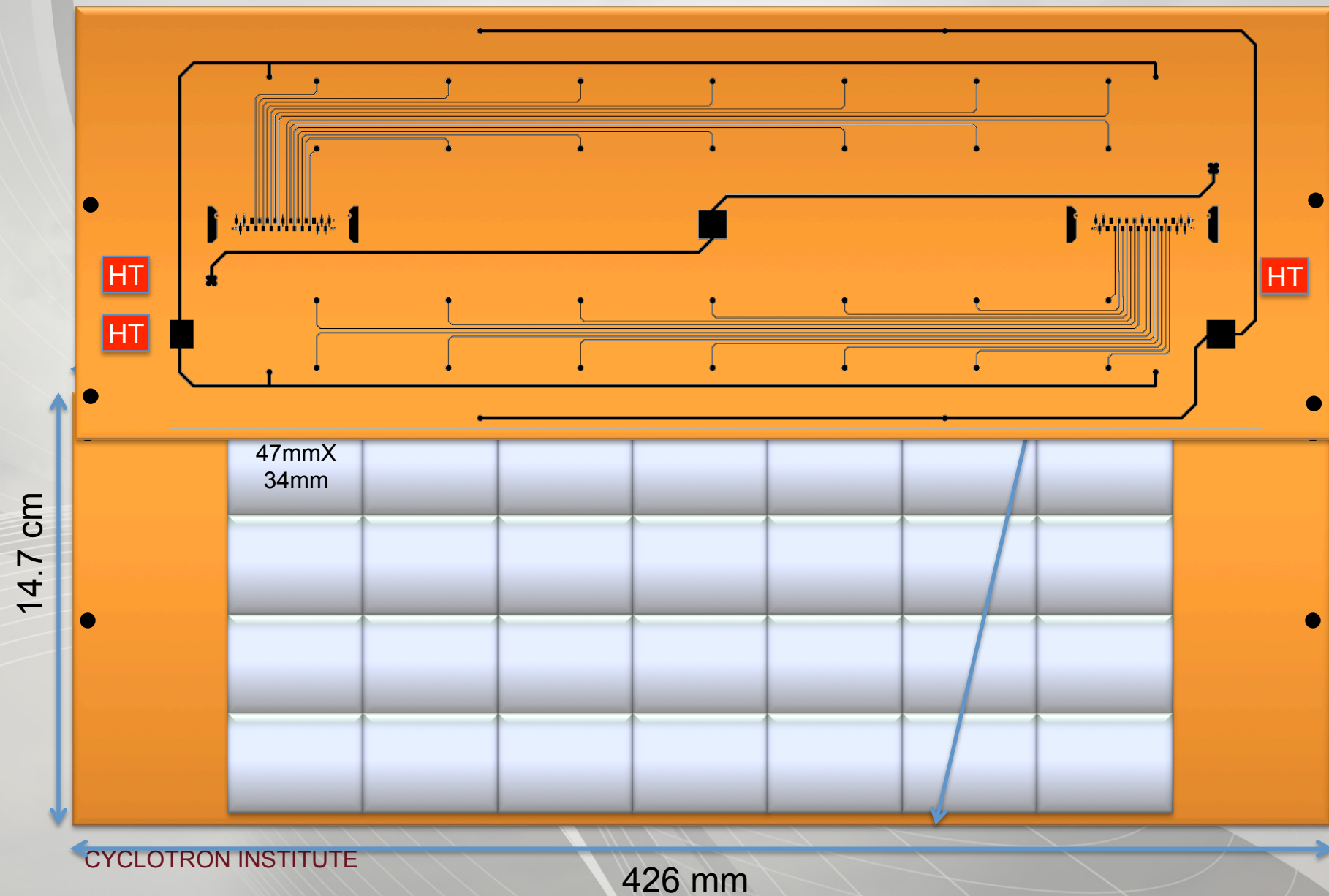
IONIZING  
PARTICLE

**Micromegas**





# New anode with Micromegas pads

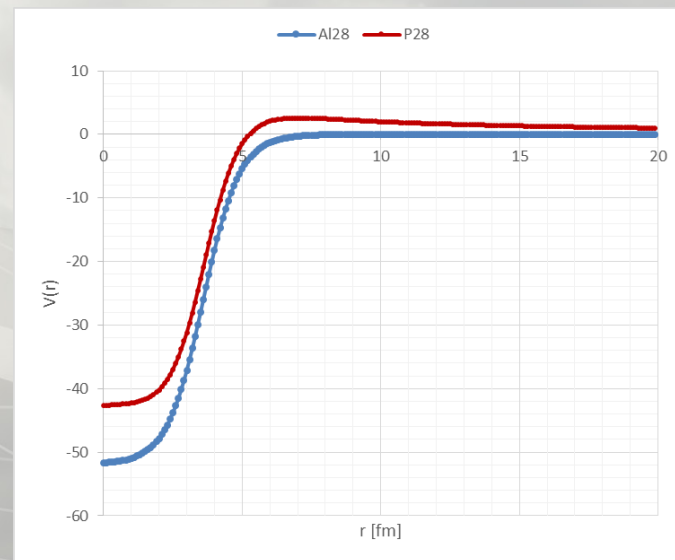
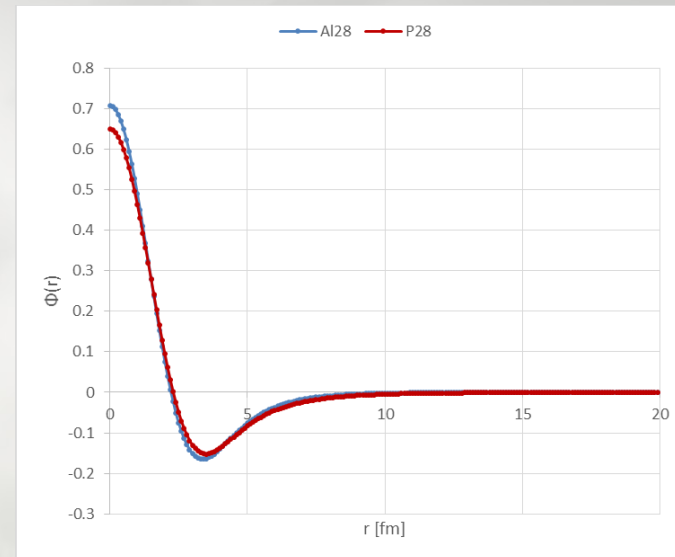


# X-Ray bursts – $^{27}\text{Si}(p,\gamma)^{28}\text{P}$

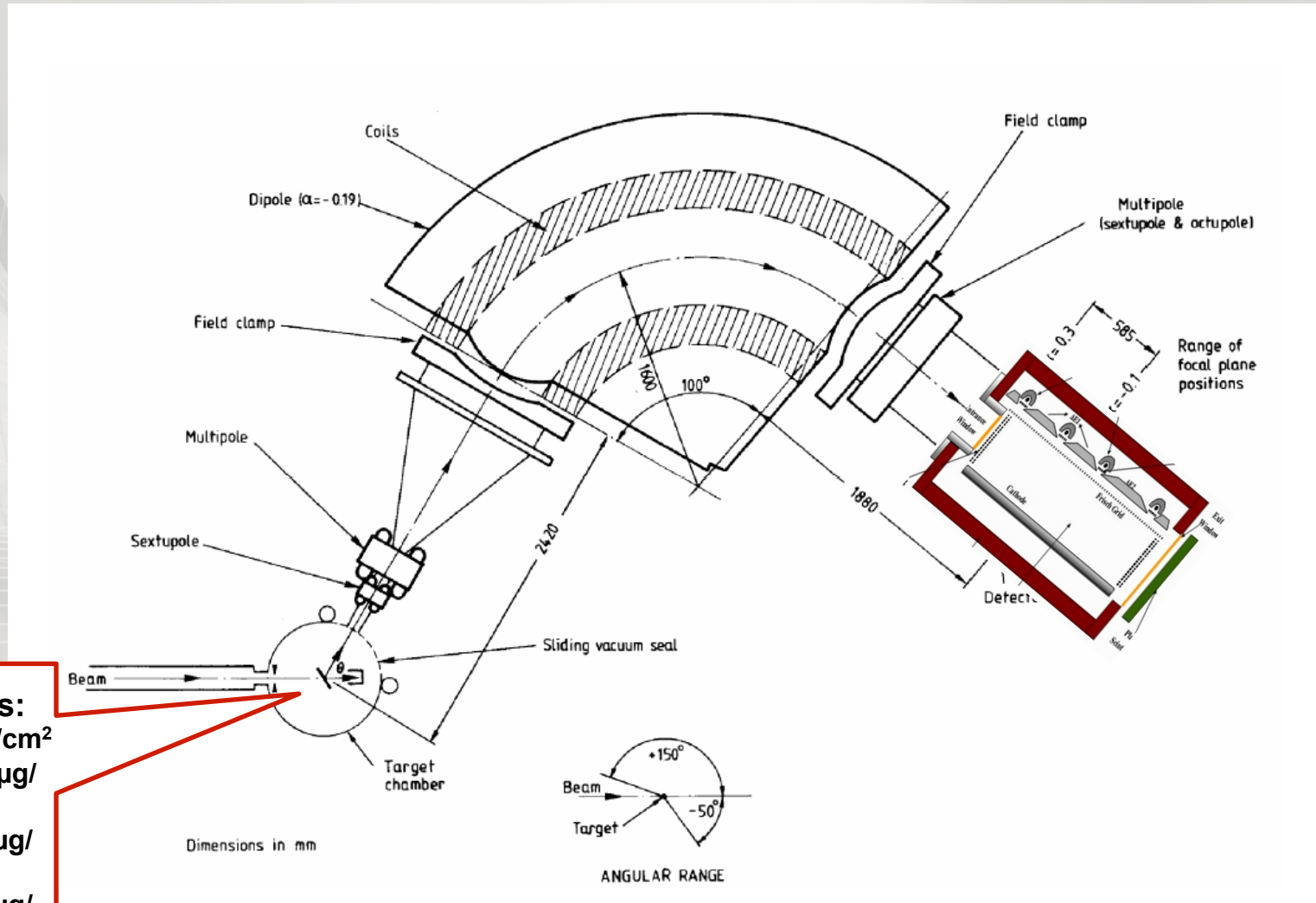
- $^{27}\text{Si}(p,g)$  – bottleneck
  - $T_{1/2}=4.15$  s ; **no experimental data**
- Use indirect method, theoretical estimations until radioactive beam available:
  - $^{28}\text{P}$  energy levels
  - Reaction Q-value
  - Spectroscopic data from mirror nucleus  $^{28}\text{Al}$

# Mirror system $^{27}\text{Al}(n, \gamma)^{28}\text{Al}$

- Study of  $^{27}\text{Al}+n \rightarrow ^{28}\text{Al}$  with MDM spectrometer
  - Beam of  $^{13}\text{C}$  @ 12 MeV/n on  $^{27}\text{Al}$  target
  - Get angular distribution from elastic scattering  $^{27}\text{Al}(^{13}\text{C}, ^{13}\text{C})^{27}\text{Al}$
  - Fit distribution to obtain Optical Potential Model parameters
  - Use OMP parameters to predict angular distribution for transfer reaction  $^{27}\text{Al}(^{13}\text{C}, ^{12}\text{C})^{28}\text{Al}$
  - Compare with experimental data to extract ANC



# TAMU Multipole-Dipole-Multipole (MDM)

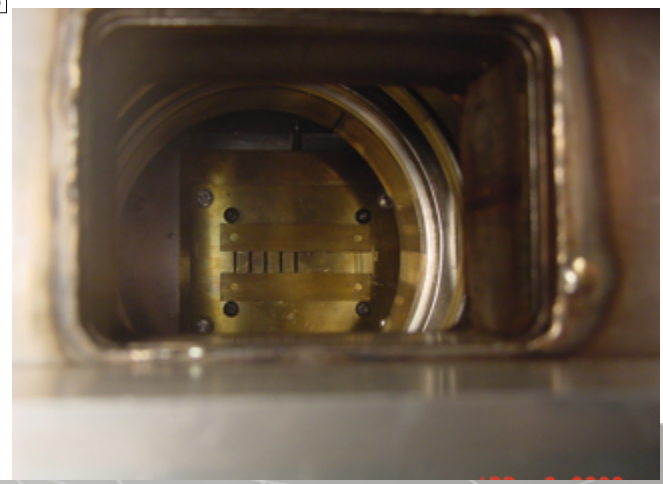
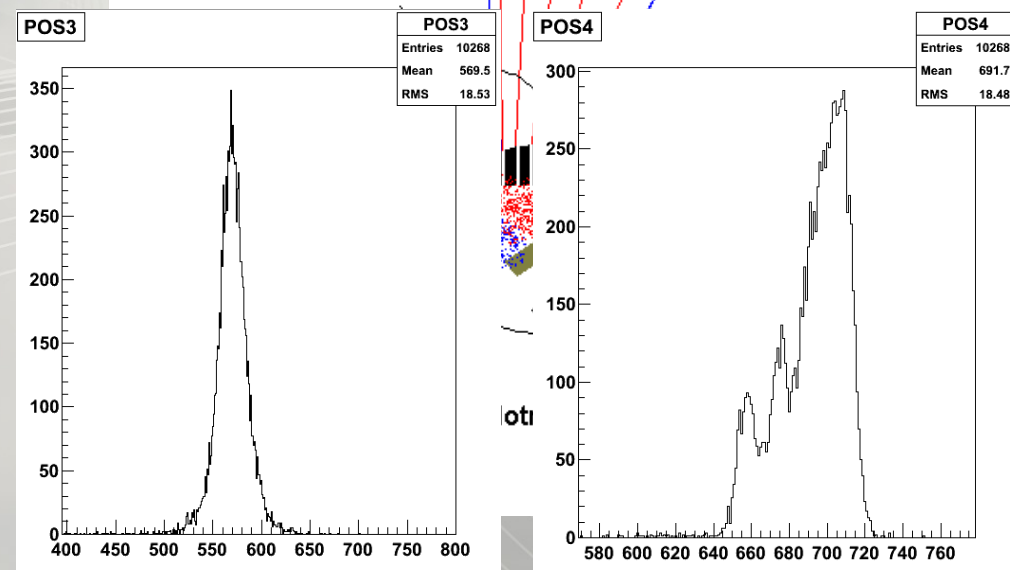
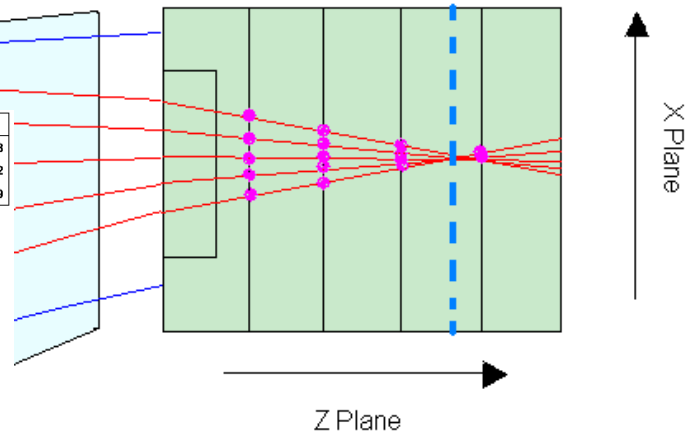
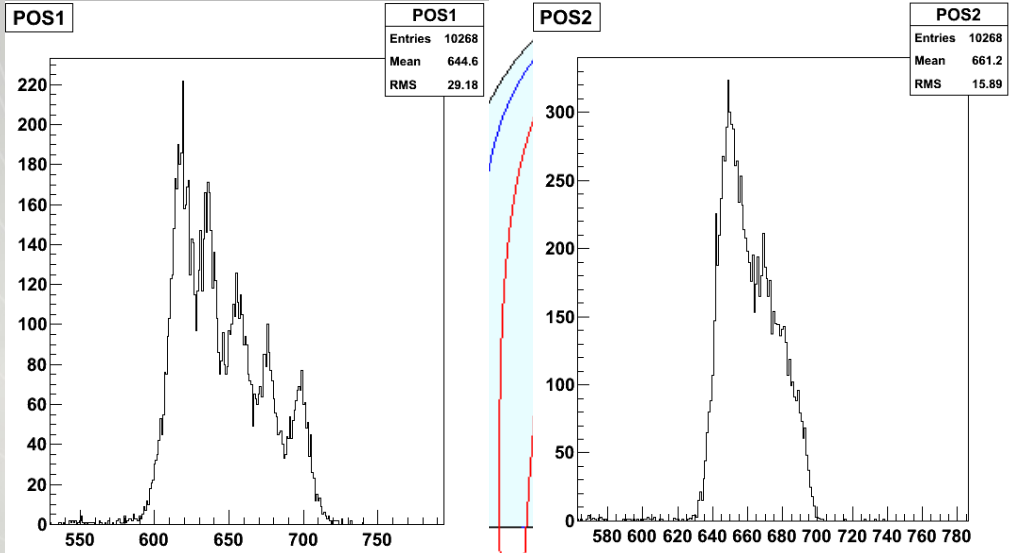


## 4 targets:

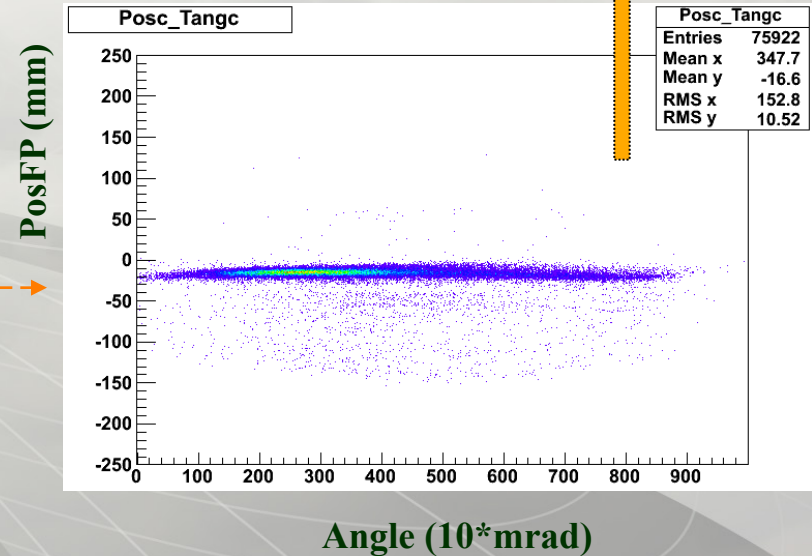
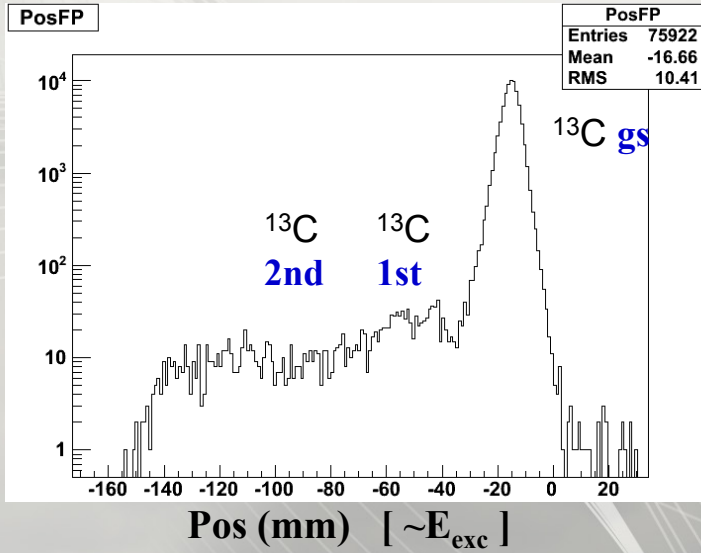
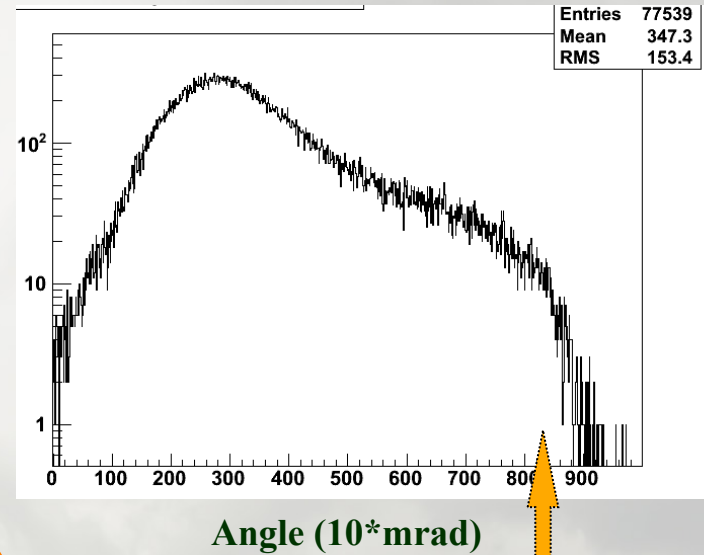
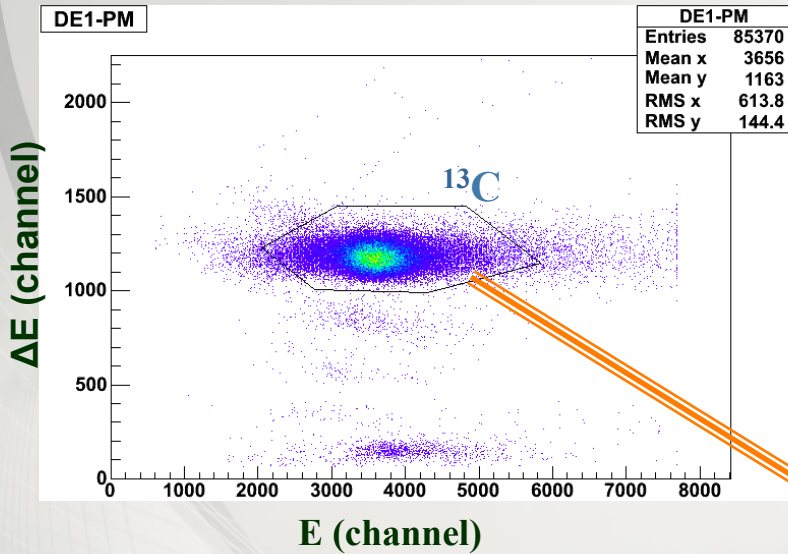
- Au 200  $\mu\text{g}/\text{cm}^2$
- <sup>27</sup>Al 105  $\mu\text{g}/\text{cm}^2$
- <sup>27</sup>Al 324  $\mu\text{g}/\text{cm}^2$
- <sup>27</sup>Al 800  $\mu\text{g}/\text{cm}^2$

# Dipole Magnet Separates Beam by Magnetic Rigidity (p/q)

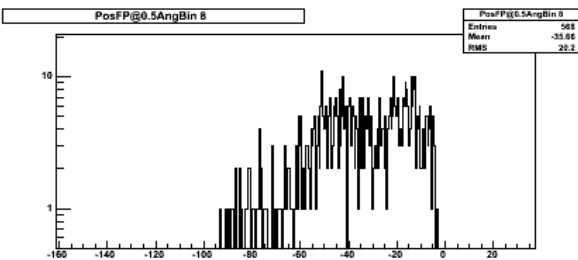
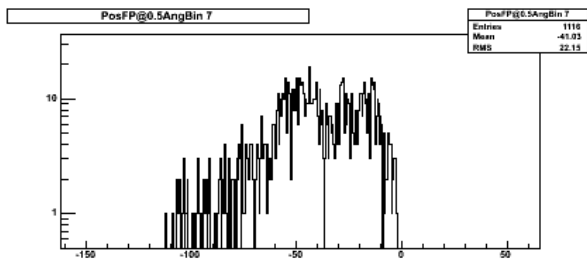
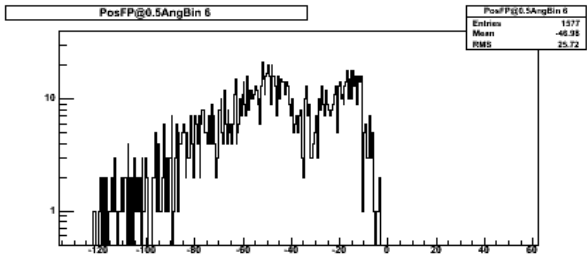
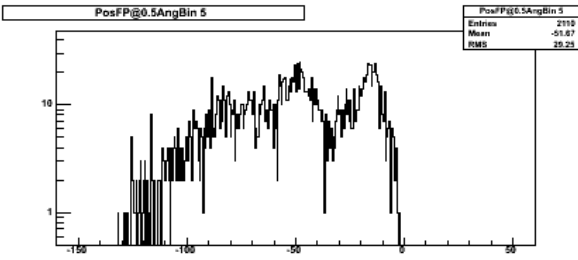
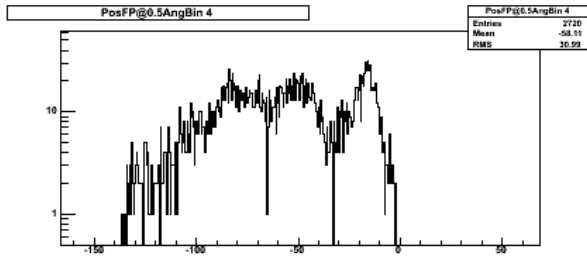
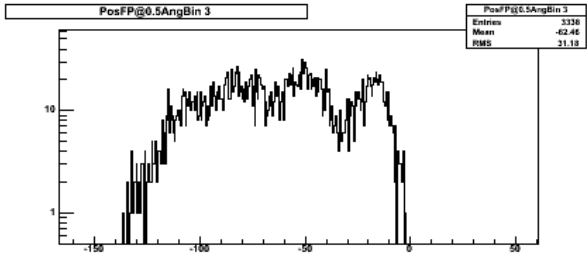
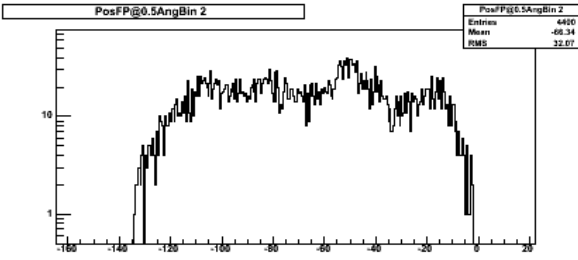
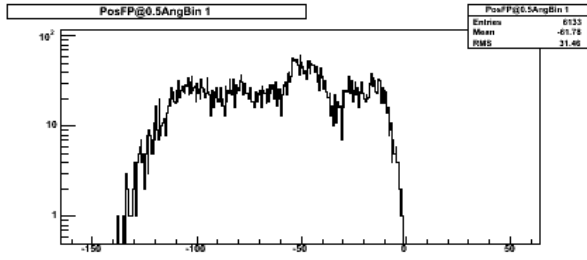
Detector



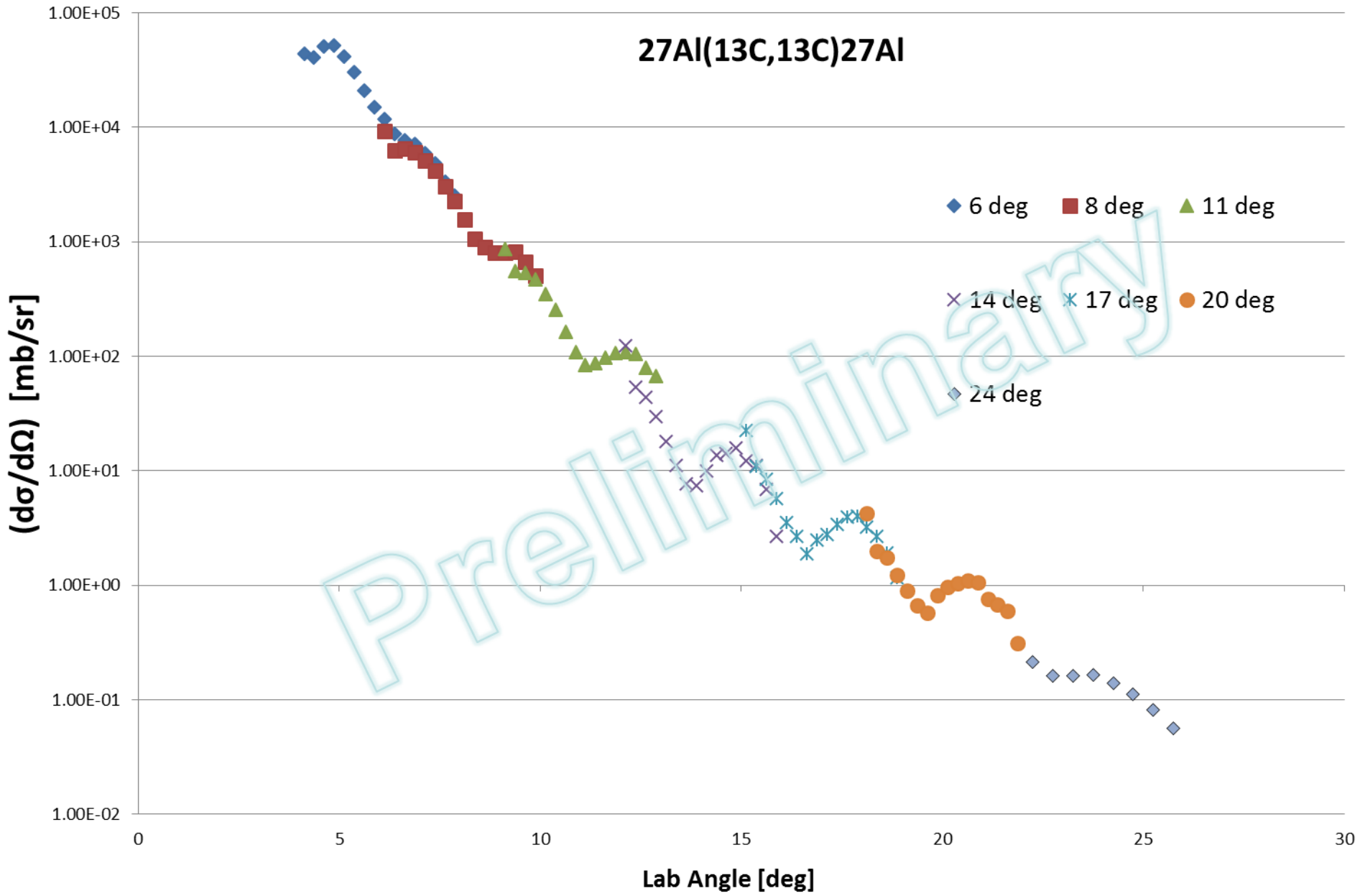
# $^{27}\text{Al}(^{13}\text{C}, ^{13}\text{C})^{27}\text{Al}$



Run 93

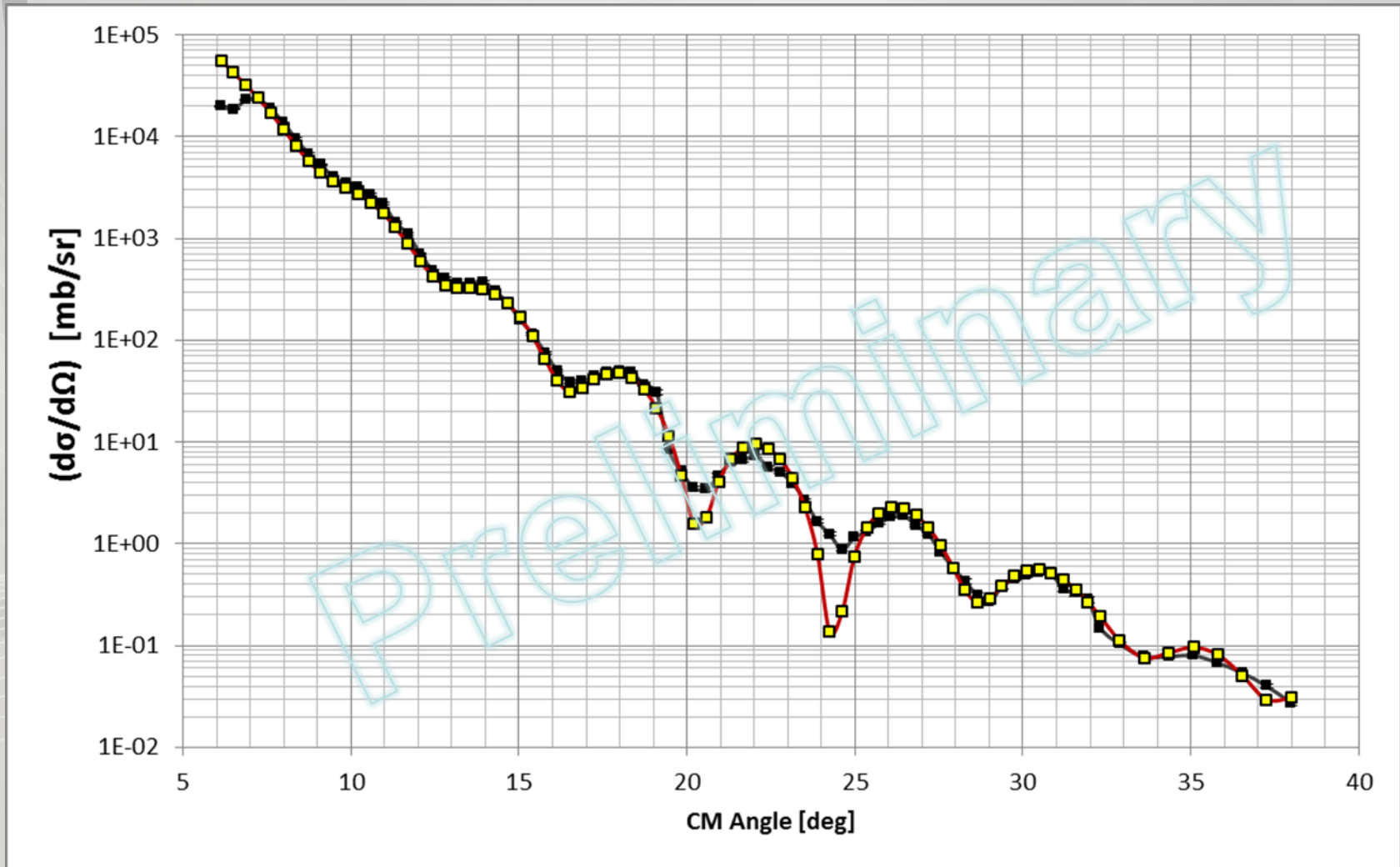


# 27Al(13C,13C)27Al



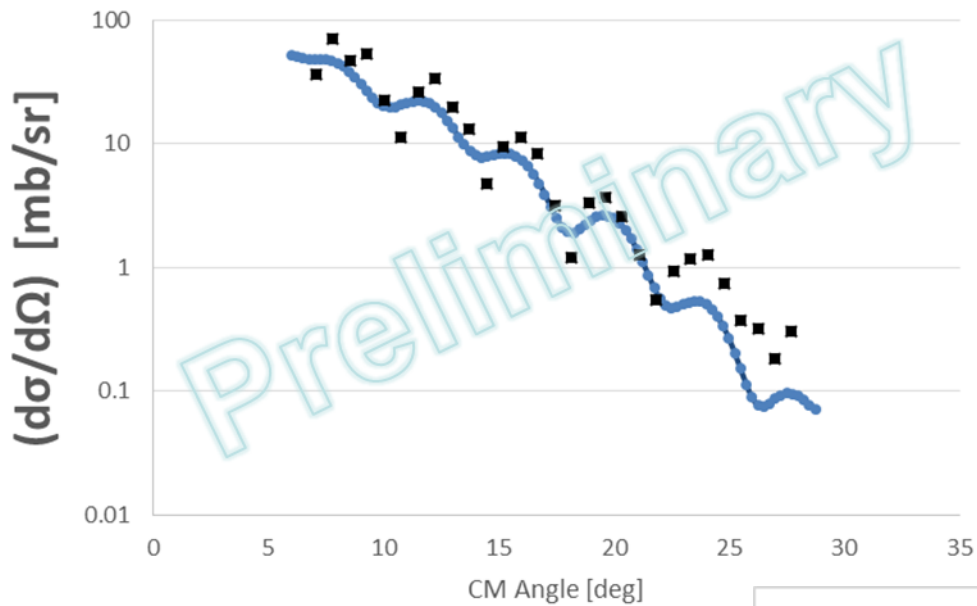


# $^{27}\text{Al}(^{13}\text{C}, ^{13}\text{C})^{27}\text{Al}$

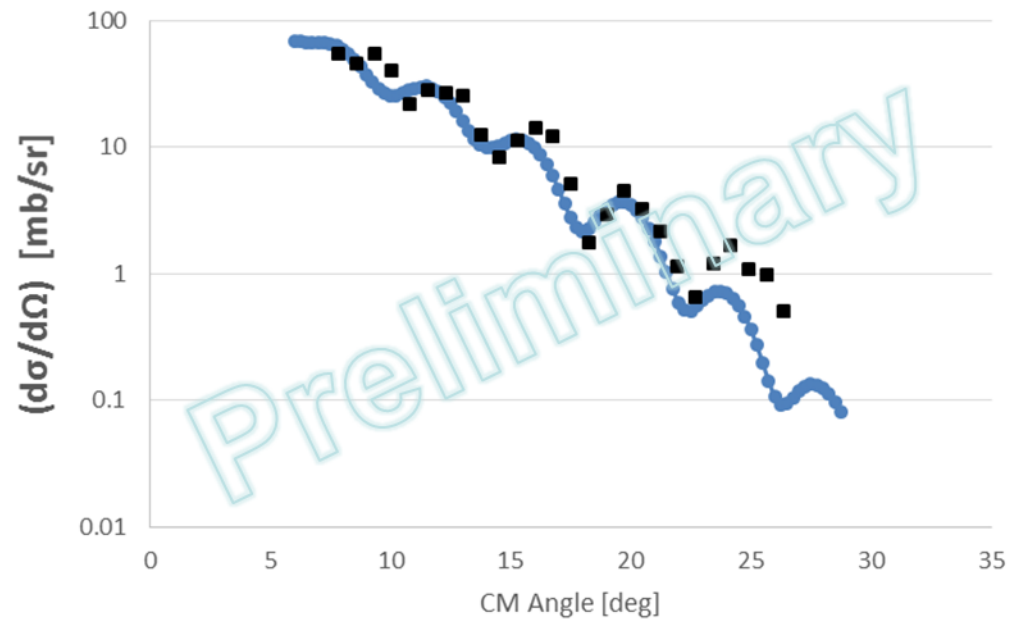


$V=71.231$   $W=14.195$   $r_V=1.046$   $r_W=1.213$   $a_V=0.537$   $a_W=0.938$   $Rc=1.764$   $\chi^2=10.17$

$^{27}\text{Al}(^{13}\text{C},^{13}\text{C})^{27}\text{Al}^*(2.212 \text{ MeV})$



$^{27}\text{Al}(^{13}\text{C},^{13}\text{C})^{27}\text{Al}^*(2.982 \text{ MeV})$



# Future plan

- Take data for  $^{27}\text{Al}(^{13}\text{C}, ^{12}\text{C})^{28}\text{Al}$  – May 2014
- Test Oxford upgrades – June 2014
- If successful, measure  $^{13}\text{C} (^{27}\text{Al}, ^{28}\text{Al}) ^{12}\text{C}$
- $^{27}\text{Si}$  beam – Fall 2014 (hopefully?)

**Thank you for your attention!**